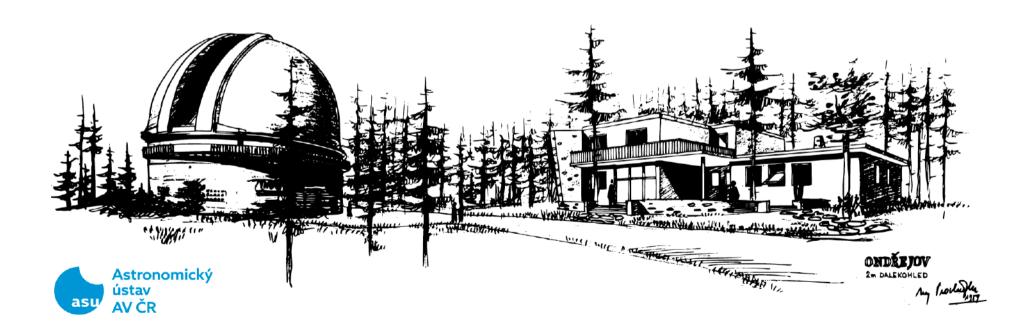
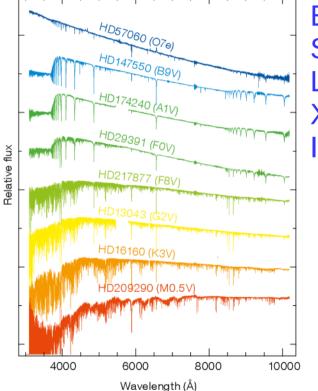
Olga Maryeva

# Reduction of spectral data from public archives



# Spectral Data Archive

# **Spectral Libraries**



Elodie Sophie LAMOST X-Shooter IACOB

# **Telescope Archives**



# SOPHIE / ELODIE on-line database of highresolution stellar spectra

## http://atlas.obs-hp.fr/elodie/

Currently **87621** (SOPHIE) + **35535** (ELODIE) spectra are fully public

ELODIE was an echelle type spectrograph installed at the Observatoire de Haute-Provence 1.93m reflector

Archive News   Publications using ELOD		The ELOD An on-line of high-resolutio	e database					Introduction   Help
Enter a designation or co	Examples: HIP117998, J04 14 5	7 15 32 10 , simbad:pr	ocyon, <u>HD190007,</u> 1	<u>HD190073, GJ</u>	<u>%1</u>			Spectra: FITS Pipeline processingDisplay the spectrumFile identification and executed pipeline: elodie:19980611/0045&z=s1d vs
b. For coordinate and around object queries, define a radius : Get spectra Get CCF Reset Query a sample of objects in a region of the sky								Coord=232845.7+520850900 Zoom out
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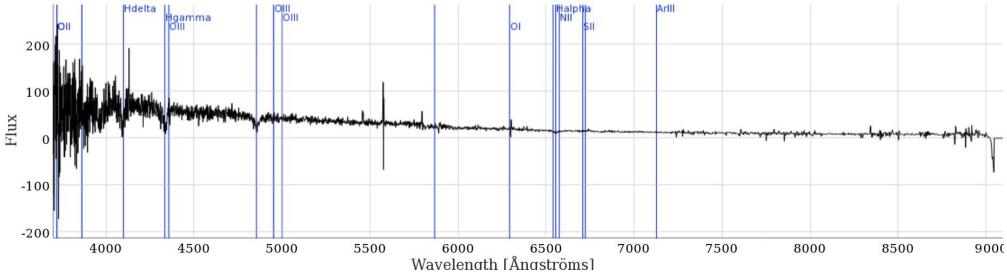
# Large Sky Area Multi-Object Fibre Spectroscopic Telescope LAMOST

### DATA RELEASE DR8 2011.10.24 - 2020.05.27

Star : **10,388,423** M Star Catalog : **773,721** A Star Catalog: **629,041** M Star Catalog : **773,721** 

### http://dr6.lamost.org/

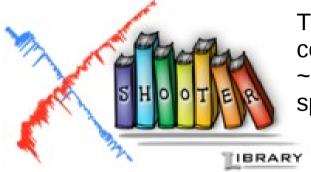




# **X-Shooter Spectral Library**

### http://xsl.u-strasbg.fr/index.html

DR2=666 stars



The X-Shooter Spectral Library is a stellar spectral library covering the wavelength range 3000–25000 Å, at a resolution R ~ 10 000. The library was built using the medium-resolution spectrograph X-Shooter (ESO, VLT).

# **ESO Library of Stellar Spectra**

http://www.eso.org/sci/facilities/paranal/decommissioned/isaac/tools/lib.html

131 stellar spectra – all spectral types

# IACOB spectroscopic database



### http://research.iac.es/proyecto/iacob/iacobcat/



Home IACOB on the sky Tecnical details Acknowledgements IACOB Publications IACOB project website 🛽

### Welcome to renewed interface of the IACOB spectroscopic database!

**IACOB** is an ambitious long-term observational project whose main scientific goal is to provide a complete and statistically significant empirical overview of the physical properties of **Galactic massive OB-type stars**. The ultimate objective of the project is that the compiled information can be used as a strong and long-lasting anchor point for our theories of stellar atmospheres, winds, interiors and evolution of massive stars.

This is the interface to have access to the spectra compiled during more than 12 years in the framework of the IACOB project: the so-called **IACOB spectroscopic database**. While not all the spectra are publicy available yet, we quote all the compiled observations for reference purposes, The different data releases (DRx) will be conveniently announced; in the meanwhile people interested in specific (samples of) spectra can contact the PI of the project by email: ssimon [at] iac.es.

More details about the project can be found in the project webpage .

The IACOB spectroscopic database mainly comprise observations made with the FIES instrument attached to the 2.56-m Nordic Optical Telescope and the HERMES spectrograph attached to the 1.2-m Mercator Telescope . In future developments, we plan to also incorporate data of Southern Galactic OB stars as obtained with the FEROS instrument (gathered from the ESO public archive), as well as long-term medium and high cadence observations of a sample of about 40 Galactic O stars and B Supergiants, collected with the 1-m Hertzsprung-SONG and the 1.2-m STELLA telescopes.

**NEWS (2020/06/21):** The second data release -- which includes 552 FIES and HERMES spectra for more than 240 late- and mid-B stars with luminosity classes V, IV and III -- is now available!.

Search by name	Spectral type	Data Release	Instrument	Only spectra available to download
Star name (Rigel, etc.)	e.g.: B1* or O* or *V*	Any	Any	Only the best spectrum per star



# Mikulski Archive for Space Telescopes – MAST

The Mikulski Archive for Space Telescopes is an astronomical data archive focused on the optical, ultraviolet, and near-infrared. MAST hosts data from over a dozen missions like Hubble, Kepler, TESS, and soon JWST.

### https://archive.stsci.edu/index.html

Home Page 🛛 💉 MAST: Schulte 7 📧											AstroView
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# Gran Telescopio CANARIAS – GTC

### http://gtc.sdc.cab.inta-csic.es/gtc/index.jsp

Object List:	Examples (coordinates in J2000): 64.99061 52.98401 4 19 57.75 +52 59 02.436 4:19:57.75 +52:59:02.436 ULAS J135058.86+081506.8
Search radius:	Browse No file selected. 5 arcmin

#### Search by Date:

Between:	16	March	~ <b>2009</b> ~
And:	26	August	~ <b>2021</b> ~

Fud Time	Exptime	A.1	Dut				Reduced Data							Raw Dat	a			
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2018-09-09 02:42:56.7	900	1.03	2	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆
2018-09-09 02:11:17.8	600	1.07	2	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆
2018-09-09 02:00:37.6	900	1.09	2	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆
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2016-08-27 02:31:12.2	600	1.11	1	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆
2016-08-27 02:20:31.7	900	1.14	1	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆
2016-08-27 02:04:58.8	1200	1.18	1	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆
2016-07-30 04:31:36.9	600	1.09	1	ADS	View	Fetch 🗆				Header	Preview	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆	View	Fetch 🗆

# SALT South African Telescope

### since July 10, 2020

# https://ssda.saao.ac.za/



Target name 😧	Resolver 😮	Proposal code 🚱	Principal investigator 🔞
	Simbad 🗸 resolve		
Right ascension 😮	Declination 😧	Proposal title 😮	Observation night 🚱
Search radius 😮	Radius units	Data category	
	Arcseconds 🐱	🗹 Science 🛛 🔲 Arcs 🔍 Bia	ses 🔲 Flats 🔲 Standard
Target type		Rejected observations	
🔲 Galaxy 🔲 ISM 🔲 Solar System Bo	ody 🔲 Star	Include rejected observation	าร

# Archive of Russian 6-m telescpe (BTA)

### https://www.sao.ru/oasis/cgi-bin/fetch?lang=en

#### The General Observation Archive Archive regulation Acknowledaments Current status **6m telescope** First date: ○ CCD 1996-02-13 - 2000-04-30 1982 ~ 01 ~ 01 ~ ○ IFP 1997-05-14 - 2000-03-02 LYNX 1996-02-28 - 2002-05-28 or MOFS 1997-03-07 - 2001-08-18 select dates by the BTA observation program key MPFS 1996-08-18 - 2009-10-27 MSS 1996-05-26 - 2017-09-11 R.A.(J2000) Decl.([2000) 1998-03-10 - 2019-10-19 (ra=hh mm ss.s; dec=[-]dd mm ss.s or in degrees) PFES 1996-08-06 - 2001-01-07 or object name: SCORPIO Radius: 5 (arcmin) SP124 1996-02-18 - 2000-12-09 **UAGS** 1994-11-08 - 2005-10-20 Data type: obs Observation mode: any ~ Small telescopes Filter: any CEGS 1997-03-24 - 2010-10-26 MAGIC 2020-01-25 - 2020-07-23 Program author: MMPP Info 2017-09-05 - 2018-12-22 MNGL ○ Z600 1996-01-12 - 2001-06-06 Clear Search ZMUAGS Radio telescope (request only by date) SCORPIO-1 is a multi-mode focal REF1 1982-01-12 - 2021-01-06 2011-10-19 - 2021-05-19 REF2 reducer of the BTA REF3 2013-02-12 - 2017-12-30 REF4 2011-07-17 - 2017-06-21 REF5 2008-09-04 - 2008-10-27 REF6 1988-02-03 - 1988-04-21



×

# Zoo of Raw Data

We have a huge variety of raw data from different instruments To properly handle them, you need to:

- Know how to work with various data formats
- Know main steps of data preparation and processing
- Know what calibration data you may need

### Main steps of data processing

- Bias / dark correction
- Cosmic rays removal
- Flat-fielding
- Mask of orders (eschelle)
- Extraction of spectra
- Linearization
- Flux calibration using spectrophotometric standard stars

### **Calibration frames**

Sometimes, calibration frames are not acquired every night, and you have to specifically look for them in data archives.

Sometimes they are part of standard observations and are distributed together with object data

# **FITS format**

De facto standard for astronomical data exchange.

Primary elements — images or binary tables

Multi-extension FITS as a way to store heterogeneous data in a single file

- Primary header with generic information
- Set of extensions with individual headers + content (image / table)
- Every extension may have human-readable name to help identifying what's inside it

Standard tools for inspection / quick look of FITS files:

**FV** — supports any FITS file with any structure, shows tables / images, plots XY graphs, etc etc

SAOImage DS9 — primarily FITS image viewer, plus rudimentary image analysis

**TopCat** — generic tool for working with data tables in various formats, including FITS tables. Supports inspection, visualization, analysis, etc

Representation of reduced spectral data in FITS

- Binary table with separate columns for wavelength and flux
- 1d/2d image + WCS standard for wavelength data
- 2d image + IRAF multispec standard for separate wavelength info for different orders

# Keywords in FITS Header (for 1D spectral data)

Representation of reduced spectral data in FITS

- Binary table with separate columns for wavelength and flux
- 1d/2d image + WCS standard for wavelength data
- 2d image + IRAF multispec standard for separate wavelength info for different orders

They all differ by FITS keywords

```
1d spectra in FITS images + WCS for linearization
```

```
Lambda = CRVAL1 + CDELT1*(x-
CRPIX1)
```

Or

```
Lambda = CRVAL1 + CD1_1*(x-CRPIX1)
```

```
X = 1..N (FITS standard is 1-based, not 0-
based!)
```

FITS standard for integer data scaling

```
Value = BZERO + data*BSCALE
e.g. 0..65535 \rightarrow -32767..32767
```

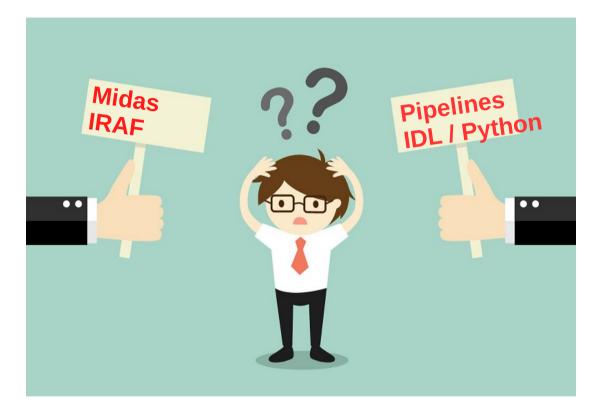
This conversion is often handled automatically by FITS reading routine, but sometimes the keywords are left in the headers — do not forget to check/remove them before saving your modified (e.g. floating-point) file!

**HISTORY** is for storing human-readable notes on the data provenance — e.g. history of data processing. Do not forget to use it in your scripts, it really helps to track various problems with the data later!

**COMMENT** is for generic comments in the header

# Data Reduction

### Midas, IRAF, Pipelines (for example ESO-reflex), IDL / Python



The choice is usually defined by your habits and existence of ready to use pipelines for specific data

- You may spend some time learning the pipeline someone else wrote
- Or spend some time writing your own processing scripts
- Or combine some existing tools you already know

No universal recipe exists, but think about re-usability of what you learned

# My experience

Togher with colleagues I wrote in IDL language a package for reduction of long-slit spectra from SCORPIO spectrograph (Russian 6-m telescope)

× 🖸		SCOrpio REduction: version 1	0	$\odot$ $\otimes$
Exit	Load log	Create log	Edit log	Help
Reduction steps: 11: Mean bias cosnic removal 22: Geometry correction 33: dividing over flat field 44: dispersion curve 55: linerization 65: spectral sensitivity normalization 17: Dispersion direction adjustment 18: Error frames calculation	Multi e Spectum	2 10 30	View input directory b), flat, neon, star, map, starslit .00, 10,00, 20,00, 30,00 10,00, 3,00 istatualit h correction using night sky lines: 0 spectrum normalization: 0 .200 y = 1 series regime: 0	View output directory Charge auxilliaries
RUN Status: II	ILe			لرا جر
I				

Later I used it to reduce the spectra from OSIRIS (GTC), FOCAS (Subaru), ISIS (WHT)

0	SPECTRUM REDUCTION	$\odot$
Load logfile	Edit logfile	Create logfile
MAIN REDUCTION STEPS:	Data Reduction Parameters:	
🔟 1 : Mean Bias	🔟 Star standart	View input directory
☐ 2: Cos mic	☐ ThAr for star standart	
☐ 3: Flat Field	Overscan values:	View output directory
☐ 4: Geometry Correction	X1: 0.00 X2: 0.00 Y1: 154.00 Y2: 172.00	View RGB
☐ 5: Create Disperse		
☐ 6: Linearization	Cosmic: OBJ FLAT THAR STAR THARSTAR	Logfile:
RUN	[10.00 [10.00 ]30.00 [10.00 ]30.00	Y
	Grating:	
	Create New Grating	
	Add New Grating	Description
SPECTRUM EXTRACTION: Extraction of spectrum of ob Extraction of spectrum of star-st	ject	
ADDITIONAL FUNCTIONS:		
Normalisation		
Barycent		
Status: IDLe		

### **Takeaway message**

- Do not be afraid of writing your own pipelines sometimes it is faster and easier than learning the code made by others (but sometimes not)
- Write for yourself, not for everyone else! It should help you, so make it as you see it fit best
- Automatize boring and repeating things
- Do not try to make too universal tools smaller and simpler tools will be easier to adapt to some new data
- Do not be afraid to express yourself in your code!

Takeaway message

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- Do not try to make too universal tools smaller and simpler tools will be easier to adapt to some new data
- Do not be afraid to express yourself in your code!

Thank you for your attention! I hope it will be useful for you