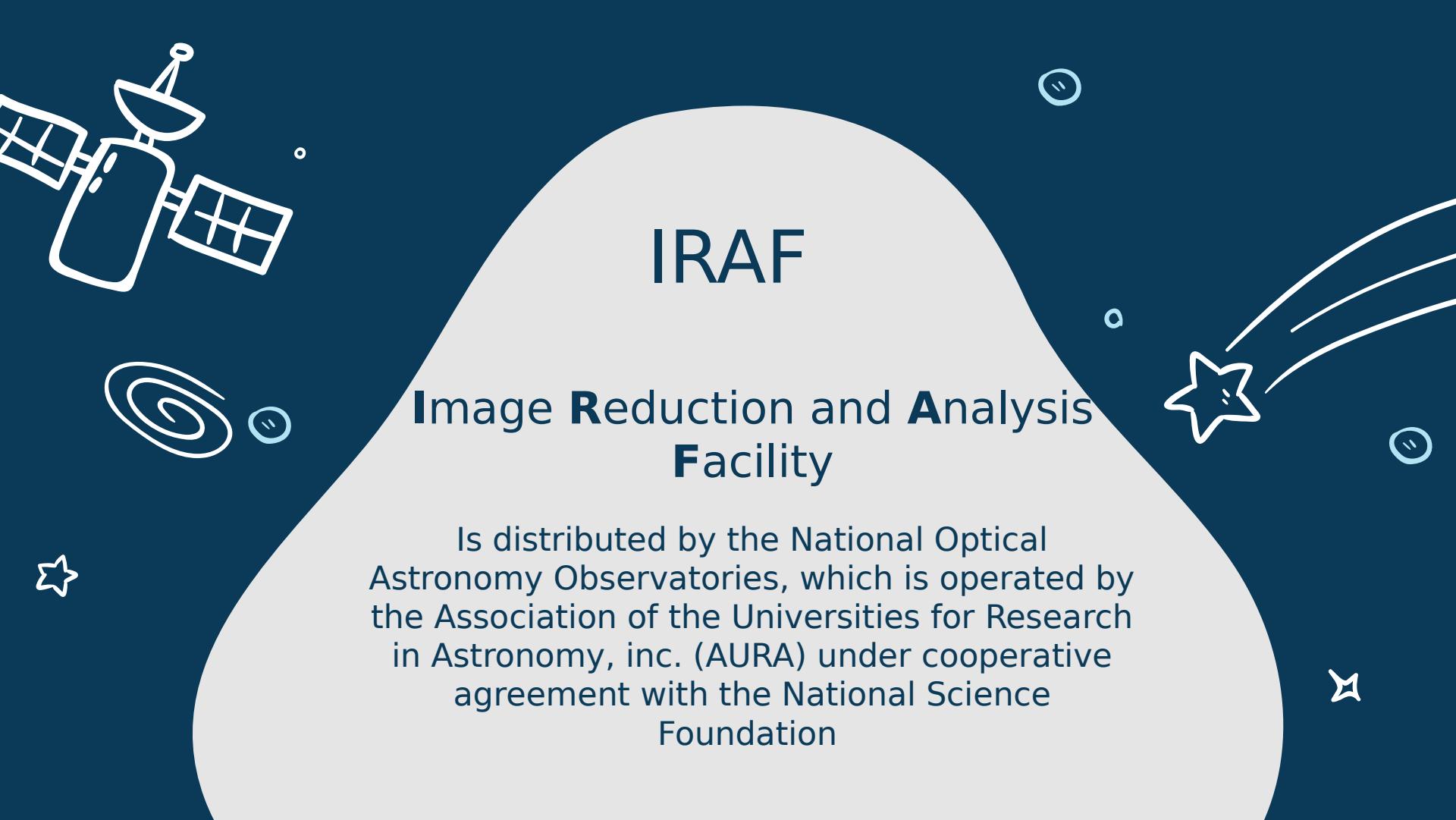


Data reduction of the echelle spectra using IRAF

Mauricio Cabezas
31.08.2022

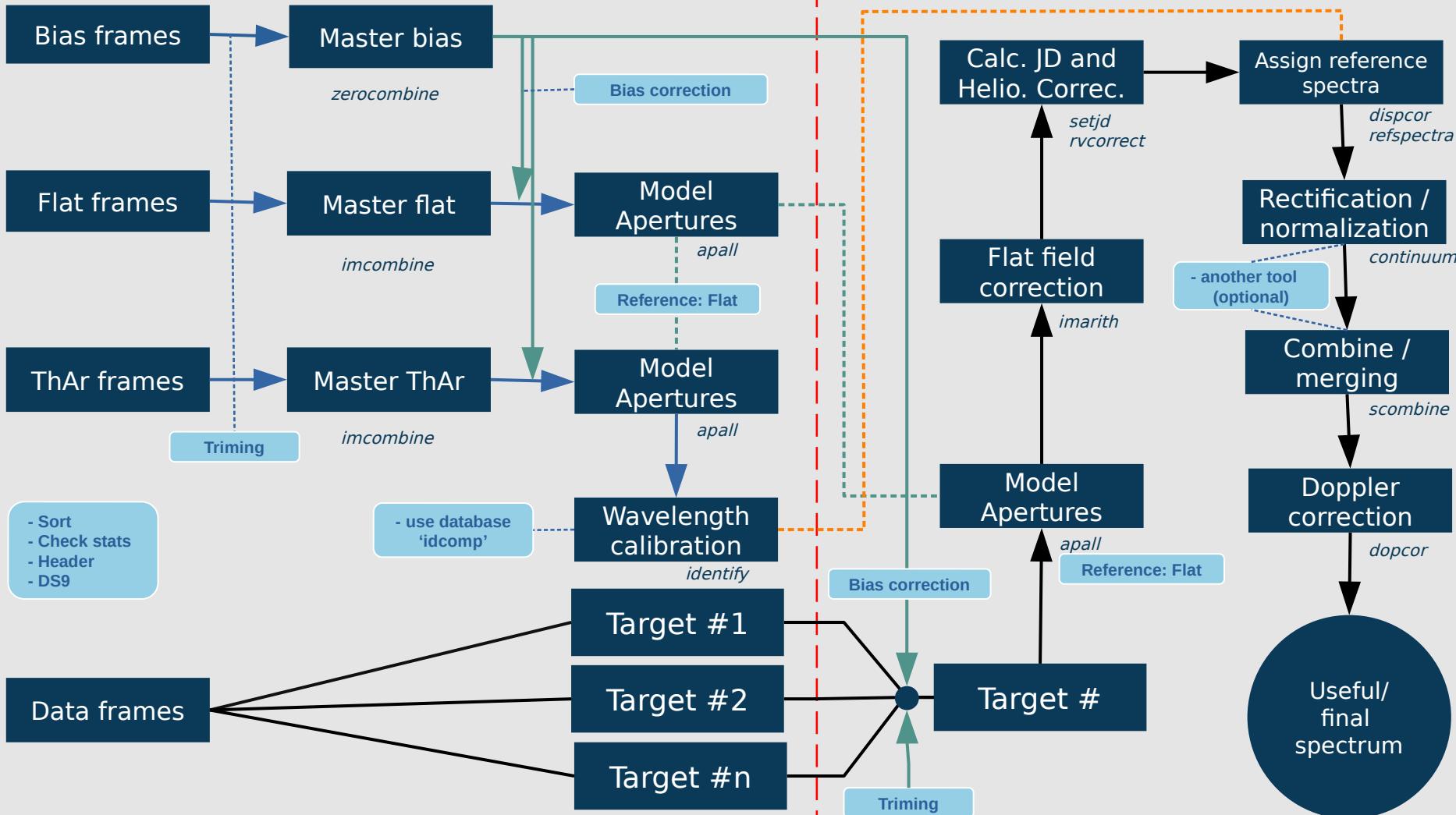
Workshop on observational techniques
29 August - 9 September 2022 at Ondřejov observatory



IRAF

Image Reduction and Analysis Facility

Is distributed by the National Optical Astronomy Observatories, which is operated by the Association of the Universities for Research in Astronomy, inc. (AURA) under cooperative agreement with the National Science Foundation



IRAF - first steps

<https://iraf.net>

help task

Plot spectrum

splot spec.fit

Plot set of spectra

specplot @spec.list

specplot e*.fit,01.fit,02.fit...

Check header

imhead spec.fit/@spec.list l+ | page

Select some field from header

hselect spec.fit/@spec.list \$l,obj-name,exptime yes

Check stats of spectrum

imstat spec.fit/@spec.list

See image with ds9

!ds9 &

display spec.fit Nframe (nframe=1,...,12)

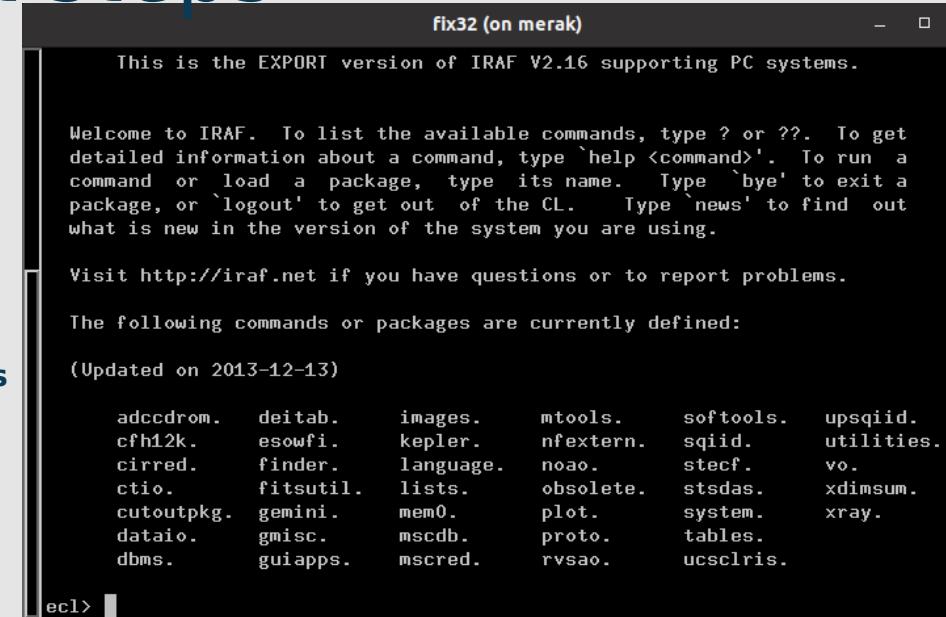
EDIT:

1) login.cl line ~34

set stdimage = imt4096

set imextn = "oif:imh fxf:fit,fits** ..."**

2) include data of Ondrejov observatory in the database, edit file **obsdb.dat** (path: ~miniconda3/envs/iraf38/iraf/noao/lib/obsdb.dat)



```
fix32 (on merak)
This is the EXPORT version of IRAF V2.16 supporting PC systems.

Welcome to IRAF. To list the available commands, type ? or ??. To get
detailed information about a command, type `help <command>'. To run a
command or load a package, type its name. Type `bye' to exit a
package, or `logout' to get out of the CL. Type `news' to find out
what is new in the version of the system you are using.

Visit http://iraf.net if you have questions or to report problems.

The following commands or packages are currently defined:

(Updated on 2013-12-13)

      adcccdrom.    deitab.    images.    mtools.    softools.    upsqiid.
      cfh12k.        esowfi.   kepler.    nfextern.   sqiid.    utilities.
      cirred.       finder.   language.   noao.     stecf.    vo.
      ctio.         fitsutil.  lists.    obsolete.   stsdas.    xdimsum.
      cutoutpkg.   gemini.    mem0.     plot.     system.    xray.
      dataio.       gmisc.    mscdb.    proto.    tables.
      dbms.        guiaapps.  mscred.   rvsao.    ucsclris.
```

ecl> |

observatory = "ondrejov"

name = "Ondrejov observatory"

longitude = 345:12:59

latitude = 49:54:38

altitude = 528

timezone = -1

OESRED.CL

task oesred=path/oesred.cl



- SEMI automatic
- always check your data!

imstat
ds9
splot

- Divided in two parts: Calibration and Science.
- Parameters were tested and works exclusively for OES.
- for now, recommend stepbystep

epar oesred (or any task- edit parameter)

Quit :q

Go! :g

The screenshot shows the IRAF command-line interface with the title 'fix32 (on merak)'. The window title is 'IRAF' and the subtitle is 'Image Reduction and Analysis Facility'. The text displayed is the help documentation for the 'oesred' task, which includes parameters for input files, calibration, object extraction, and normalization.

```
PACKAGE = clpackage
TASK = oesred

input   = e202102040008.fit Spectrum target to reduce(.fit)
(output = hd54482) Output filename
(idtarge= HD 54482) Target name on header
(napertu= 49) Number of apertures to be found
(id     = 0008) Observation id number

# CALIBRATION PARAMETERS

(orgfile= no) do you want organize files?
(zerocom= no) Combine zero level images?
(trimcal= no) Trim flat and comp?
(iftrimc= no) Use trim flat & comp?
(zerocor= no) Apply zero level correction to flat & comp?
(compcom= no) Combine comparison lamp images?
(flatcom= no) Combine flat field images?
(flatapa= no) Extract flat apertures?
(compapa= no) Extract comparison apertures?
(iddatab= no) Use database folder for identification?
(idfolde= idcomp) folder name with identification database
(idencom= no) Identify features in spectrum for dispersion sol

# OBJECT PARAMETERS

(trimob = no) Trim object?
(iftrimo= no) Use trim object?
(zerocor= no) Apply zero level correction to object?
(crays  = no) Remove cosmic rays?
(ifcrays= no) Use object with cosmic rays extraction?
(objecta= no) Extract object apertures?
(flatcor= no) Apply flat correction to object?
(helioco= no) calculate JD + heliocentric correction?
(idref  = no) refer database identification to images?
(norm   = no) normalize spectra?
(ncombin= no) combine normalized spcectra?
(mode   = ql)
```

1) Initial parameters!

Check type of "image":

flat: flat fields

zero: bias

comp: ThAr comparison spectra

Object: science

imhead e*.fit (short way)

- print list filename and object type

hselect e*.fit \$l,object yes

- check header!

imhead e202109060001.fit l+ / page

- for example, we wanna start with alp Lyr.

- take note about the filename and the target name in the header!

```
Image Reduction and Analysis Facility
PACKAGE = clpackage
TASK = oesred

input   =   e202109060016.fit Spectrum target to reduce(.fit)
(output =           alpLyr) Output filename
(idtarg=           alp Lyr) Target name on header
(napertu=          49) Number of apertures to be found
(id     =          0016) Observation id number
```

```
fix32 (on merak)

SGH-OIC = 2 / OES Iodine cell
TM-DIFF = -1 / T1630950141 - P1630950142
OBJECT = 'zero' / Title of observation
IMAGETYP= 'zero' / Type of observation, eg. FLAT
OBSERVER= 'Geier Rezba' / Observers
SYSVER = 'PESO exported.exp'
READSPD = '100kHz'
FILENAME= 'e202109060001.fit'
CAMFOCUS= 3080. / Camera focus position
SPECTEMP= 21.8 / 17912
SPECFILT= 0 / Spectral filter
SLITHEIG= 1.07 / Slit hight in mm
TM_START= 63741 / 17:42:21, 1630950141
UT      = '17:42:21' / UTC of start of observation
EPOCH   = 2000. / Same as EQUINOX - for back compat
EQUINOX = 2000. / Equinox of RA and DEC
DATE-OBS= '2021-09-06' / UTC date start of observation
TM_END  = 63742 / 17:42:22, 1630950142
EXPTIME = 1 / Length of observation excluding pauses
DARKTIME= 1 / Length of observation including pauses
CCDTEMP = -110 / Detector temperature
STDIN-line 63-file 1 of 1
```

e202109060013.fit	comp
e202109060014.fit	comp
e202109060015.fit	comp
e202109060016.fit	"alp Lyr"
e202109060017.fit	"alp Lyr"
e202109060018.fit	"HD 10780"
e202109060019.fit	"BD+44 2417"
e202109060020.fit	"HD 153911"
e202109060021.fit	"KIC 11134982"
e202109060022.fit	"HD 209027"
e202109060023.fit	"HD 194905"
e202109060024.fit	comp
e202109060025.fit	comp
e202109060026.fit	comp

ALWAYS keep raw data as backup!!

```
          Image Reduction and Analysis Facility
PACKAGE = clpackage
         TASK = oesred

input   =    e202109060016.fit  Spectrum target to reduce(.fit)
(output =                  alpLyr) Output filename
(idtarge=                  alp Lyr) Target name on header
(napertu=                  49) Number of apertures to be found
(id      =                  0016) Observation id number
```

input: The complete fit file name of our target, science

output: Filename of the final reduced spectrum > “output-id.fit”

Idtarget: EXACT name of the target which is in the header!!

naperture: Number of apertures, important if we wanna use the wavelength calibration database.

id: recommended the last four numbers of the fit file

2) sort files!

- complete first set of parameters, in the part of calibration set:

```
orgfile = yes  
:go
```

After this step files are organized, you can check each folder and files.

Example:

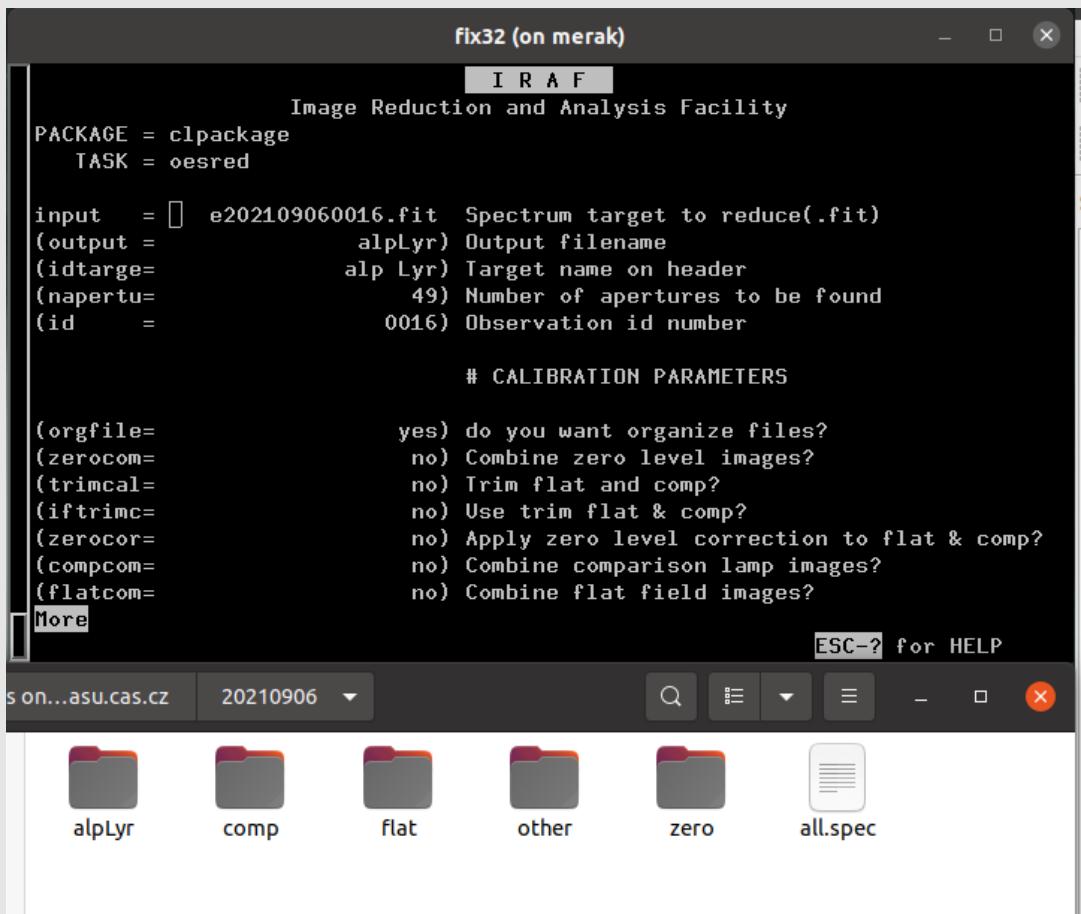
```
imstat @flat.dat  
# IMAGE NPIX MEAN STDDEV MIN MAX  
e202109060006.fit 4194304 931.1 2621. 0. 25203.  
e202109060007.fit 4194304 949.1 2679. 0. 25797.  
e202109060008.fit 4194304 944.7 2664. 0. 25560.  
e202109060009.fit 4194304 941.4 2654. 0. 25521.  
e202109060010.fit 4194304 940.2 2649. 0. 25405.
```

- visual inspection:

```
!ds9 &  
display e202109060006.fit 1
```

Before each run be
sure we are in the
main folder!

```
pwd
```



3) Combine zeros!

- before, is possible edit the list **zero.dat**
- change the previous task to "no"

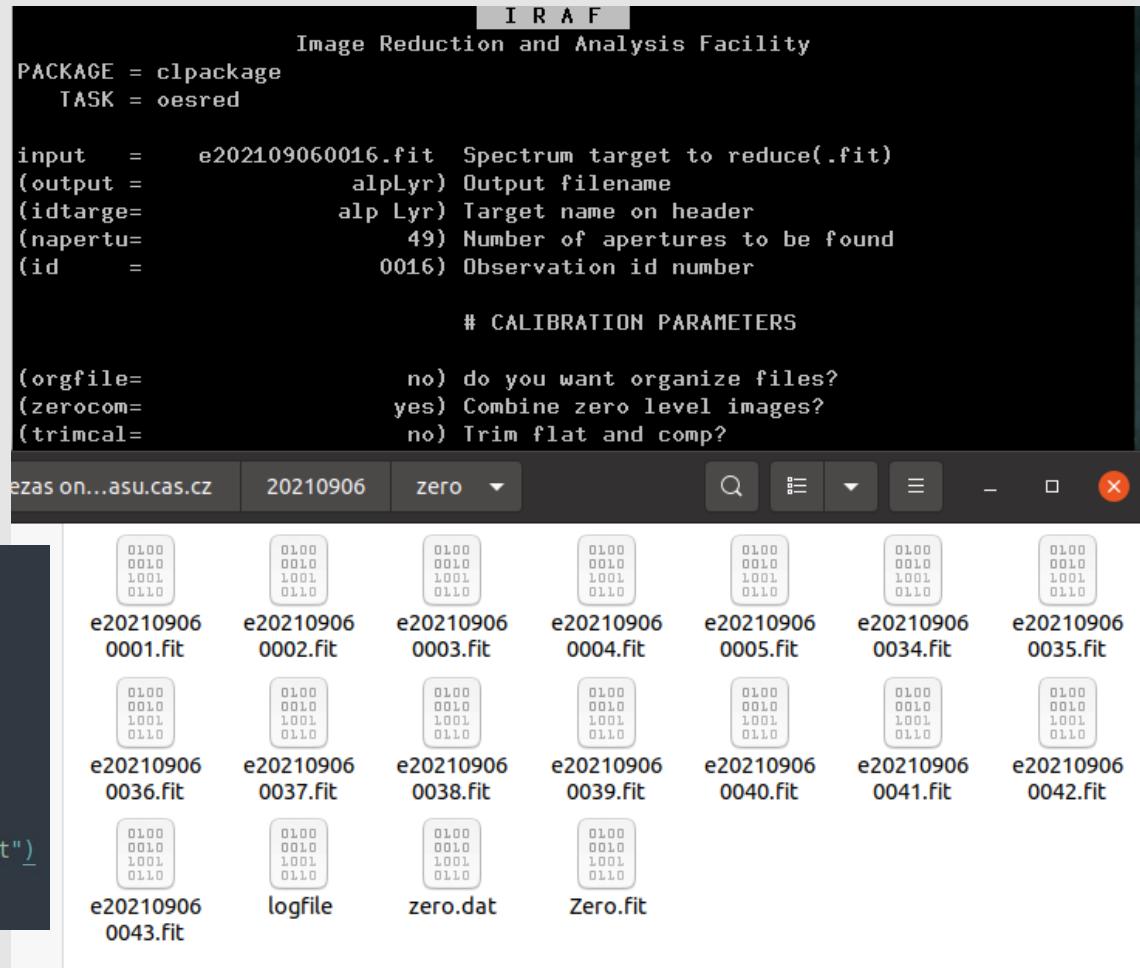
orgfile = no

zerocom = yes

:go

- new file: **Zero.fit**

```
104 ##### ZEROCOMBINE
105 if (access("zero/Zero.fit")){
106     zerocomb=no
107 }
108 if (zerocomb==yes){
109     cd "zero/"
110     unlearn zerocombine
111     zerocombine.reject="minmax"
112     zerocombine.rdnnoise= "READNOIS"
113     zerocombine.gain  = "GAIN"
114     zerocombine (input="@zero.dat",output="Zero.fit")
115     cd "../"
116 }
```



4) trim flat and comp (optional)

- completely optional, but sometimes the first and last pixel of each aperture is saturated and can be annoying during normalization.

- lines 120 in oesred.cl

zeroicom = no

trimcal = yes

iftrimc = yes

:go

- new files, prefix **T**

Te*.fit

```
IRAF
Image Reduction and Analysis Facility

PACKAGE = clpackage
TASK = oesred

input   = e202109060016.fit Spectrum target to reduce(.fit)
(output = alplyr) Output filename
(idtarge= alp Lyr) Target name on header
(napertu= 49) Number of apertures to be found
(id      = 0016) Observation id number

# CALIBRATION PARAMETERS

(orgfile= no) do you want organize files?
(zeroocom= no) Combine zero level images?
(trimcal= yes) Trim flat and comp?
(iftrimc= yes) Use trim flat & comp?
(zeroocor= no) Apply zero level correction to flat & comp?
(compcom= no) Combine comparison lamp images?
(flatcom= no) Combine flat field images?
(flatapa= no) Extract flat apertures?
(compapa= no) Extract comparison apertures?

More
ecl> 
```

bezas on...asu.cas.cz 20210906 flat ▾

	e20210906_0006.fit	e20210906_0007.fit	e20210906_0008.fit	e20210906_0009.fit	e20210906_0010.fit	flat.dat	logfile
	e20210906_0006.fit	e20210906_0007.fit	e20210906_0008.fit	e20210906_0009.fit	e20210906_0010.fit	flat.dat	logfile
	Te20210906_0006.fit	Te20210906_0007.fit	Te20210906_0008.fit	Te20210906_0009.fit	Te20210906_0010.fit		

4) trim flat and comp (optional)

If we apply trim and you wanna use the trimmed images, you need to set iftrimc=yes during the whole reduction!

```
125  if (trimcal==yes){  
126      unlearn ccdproc  
127      ccdproc.trimsec = "[2:2035,*]"  
128      # ccdproc.trimsec = "[5:2025,800:1500]"  
129      ccdproc.trim = yes  
130      ccdproc.fixpix = no  
131      ccdproc.overscan = no  
132      ccdproc.darkcor= no  
133      ccdproc.zerocor=no  
134      ccdproc.flatcor=no  
135      #  
136      cd "flat/"  
137      ccdproc.ccdtype = "flat"  
138      ccdproc (images="@flat.dat",output="T@flat.dat")  
139      cd "../"  
140      #  
141      cd "comp/"  
142      ccdproc.ccdtype = "comp"  
143      ccdproc (images="@comp.dat",output="T@comp.dat")  
144      cd "../"  
145  
146  }  
147
```

5) Bias correction -> flat and comp

trimcal = no

iftrimc = yes

zerocor = yes

:go

- new files, prefix **Z**

ZTe*.fit

```
I R A F
Image Reduction and Analysis Facility

PACKAGE = clpackage
TASK = oesred

input   = e202109060016.fit Spectrum target to reduce(.fit)
(output = alpLyr) Output filename
(idtarg= alp Lyr) Target name on header
(napertu= 49) Number of apertures to be found
(id     = 0016) Observation id number

# CALIBRATION PARAMETERS

(orgfile= no) do you want organize files?
(zerocom= no) Combine zero level images?
(trimcal= no) Trim flat and comp?
(iftrimc= yes) Use trim flat & comp?
(zerocor= yes) Apply zero level correction to flat & comp?
(compcom= no) Combine comparison lamp images?
(flatcom= no) Combine flat field images?
(flatapa= no) Extract flat apertures?
(compapa= no) Extract comparison apertures?

More
ecl>
```

bezas on...asu.cas.cz	20210906	flat	Q	E	▼	☰	×
	e20210906 0006.fit	e20210906 0007.fit	e20210906 0008.fit	e20210906 0009.fit	e20210906 0010.fit	flat.dat	logfile
	Te20210906 0006.fit	Te20210906 0007.fit	Te20210906 0008.fit	Te20210906 0009.fit	Te20210906 0010.fit	ZTe2021090 60006.fit	ZTe2021090 60007.fit
	ZTe2021090 60008.fit	ZTe2021090 60009.fit	ZTe2021090 60010.fit				

5) Bias correction flat and comp

```
149 if (zerocorcal==yes){
150     unlearn ccdproc
151     ccdproc.ccdtype="zero"
152     ccdproc.fixpix = no
153     ccdproc.overscan = no
154     ccdproc.darkcor= no
155     ccdproc.zerocor=no
156     ccdproc.flatcor=no
157     #
158     cd "flat/"
159     ccdproc.ccdtype="flat"
160     ccdproc.zerocor=yes
161     ccdproc.zero="../zero/Zero.fit"
162     if (iftrimc==yes){
163         ccdproc (images="T@flat.dat",output="ZT@flat.dat")
164     } else {
165         ccdproc (images="@flat.dat",output="Z@flat.dat")
166     }
167     cd "../"
168     #
169     cd "comp/"
170     ccdproc.ccdtype="comp"
171     ccdproc.zerocor=yes
172     ccdproc.zero="../zero/Zero.fit"
173         if (iftrimc==yes){
174             ccdproc (images="T@comp.dat",output="ZT@comp.dat")
175         } else {
176             ccdproc (images="@comp.dat",output="Z@comp.dat")
177         }
178     cd "../"
179 }
```

6) Combine comp (ThAr)

iftrimc = yes (keep "yes")

zerocor = no

compcom = yes

:go

- new file

ZTcomp.fit

```
IRAF
Image Reduction and Analysis Facility

PACKAGE = clpackage
TASK = oesred

input   = e202109060016.fit Spectrum target to reduce(.fit)
(output = alpLyr) Output filename
(idtarge= alp Lyr) Target name on header
(napertu= 49) Number of apertures to be found
(id      = 0016) Observation id number

# CALIBRATION PARAMETERS

(orgfile= no) do you want organize files?
(zeroicom= no) Combine zero level images?
(trimcal= no) Trim flat and comp?
(iftrimc= yes) Use trim flat & comp?
(zerocor= no) Apply zero level correction to flat & comp?
(compcom= yes) Combine comparison lamp images?
(flatcom= no) Combine flat field images?
(flatapa= no) Extract flat apertures?
(compapa= no) Extract comparison apertures?

More
ESC-? for HELP

ezas on...asu.cas.cz 20210906 ▾ Q ☰ - ☐ ×

alpLyr comp flat other zero all.spec ZTcomp.fit
```

7) Combine flat field

iftrimc = yes

compcom = no

flatcom = yes

:go

- new file

ZTflat.fit

```
I R A F
Image Reduction and Analysis Facility
PACKAGE = clpackage
  TASK = oesred

input   = e202109060016.fit Spectrum target to reduce(.fit)
(output = alpLyr) Output filename
(idtarg= alp Lyr) Target name on header
(napertu= 49) Number of apertures to be found
(id     = 0016) Observation id number

# CALIBRATION PARAMETERS

(orgfile= no) do you want organize files?
(zerocom= no) Combine zero level images?
(trimcal= no) Trim flat and comp?
(iftrimc= yes) Use trim flat & comp?
(zerocon= no) Apply zero level correction to flat & comp?
(compcom= no) Combine comparison lamp images?
(flatcom= yes) Combine flat field images?
(flatapa= no) Extract flat apertures?
(compapa= no) Extract comparison apertures?

More
ESC-? for HELP

ezas on...asu.cas.cz 20210906 ▾
Q   E   ▾   ▾
alpLyr  comp  flat  other  zero  all.spec  ZTcomp.fit
0100
0010
1001
0110
ZTflat.fit
```

6) and 7) Combine comp/flat field

```
182 ##### COMBINE - comp/lamp
183 if (compcomb==yes){
184     cd "comp/"
185     unlearn imcombine
186     imcombine.reject = "none"
187     imcombine.lsigma = 3
188     imcombine.hsigma = 3
189     imcombine.rdnnoise= "READNOIS"
190     imcombine.gain    = "GAIN"
191     imcombine.scale   = "exposure"
192     imcombine.expname="EXPTIME"
193 #      imcombine (input="@comp.dat",output = "comp.fits")
194 #          if (iftrimc==yes){
195 #              imcombine (input="ZT@comp.dat",output = "../ZTcomp.fit")
196 #          } else {
197 #              imcombine (input="Z@comp.dat",output = "../Zcomp.fit")
198 #          }
199     cd "../"
200 }
```

8) Model apertures - Flat

iftrimc = yes

flatcom = no

flatapa = yes

:go

Find apertures for Ztflat? (yes):

Number of apertures to be found automatically (49):

Resize apertures for ZTflat? (yes):

Edit apertures for Ztflat? (yes):

- accept everything with “enter” or typing “yes”. Default answer in parenthesis.

- in order to do everything a bit easy/fast. We will use a database for wavelength calibration, for that we need to choose 49 apertures.

Background fitting is not necessary because overlapping.

```
TASK = oesred
input   = e202109060016.fit Spectrum target to reduce(.fit)
(output  = alpLyr) Output filename
(idtarg= alp Lyr) Target name on header
(napertu= 49) Number of apertures to be found
(id      = 0016) Observation id number

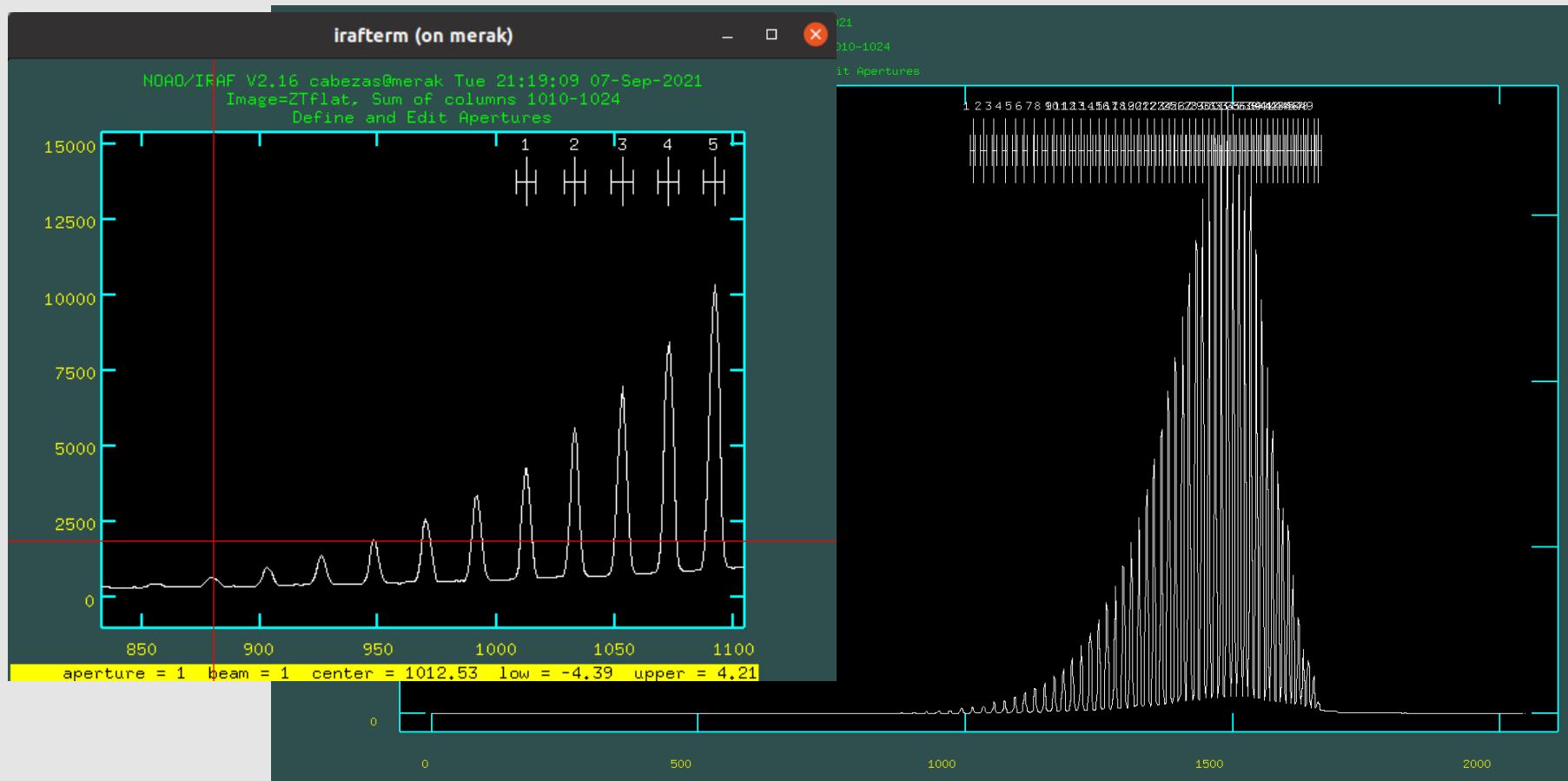
# CALIBRATION PARAMETERS

(orgfile= no) do you want organize files?
(zerocom= no) Combine zero level images?
(trimcal= no) Trim flat and comp?
(iftrimc= yes) Use trim flat & comp?
(zerocor= no) Apply zero level correction to flat & comp?
(compcom= no) Combine comparison lamp images?
(flatcom= no) Combine flat field images?
(flatapa= yes) Extract flat apertures?
(compapa= no) Extract comparison apertures?

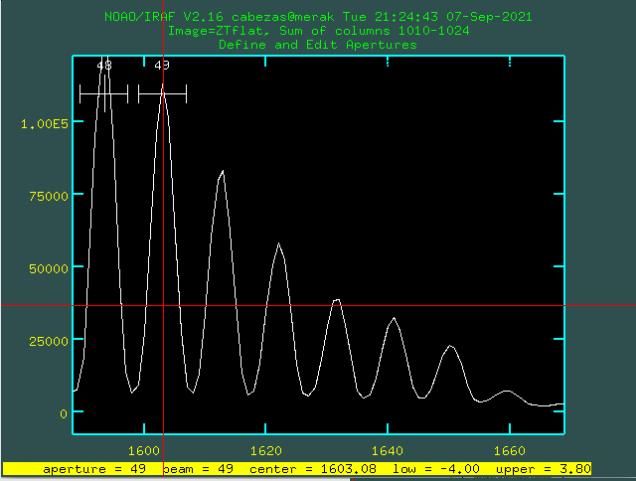
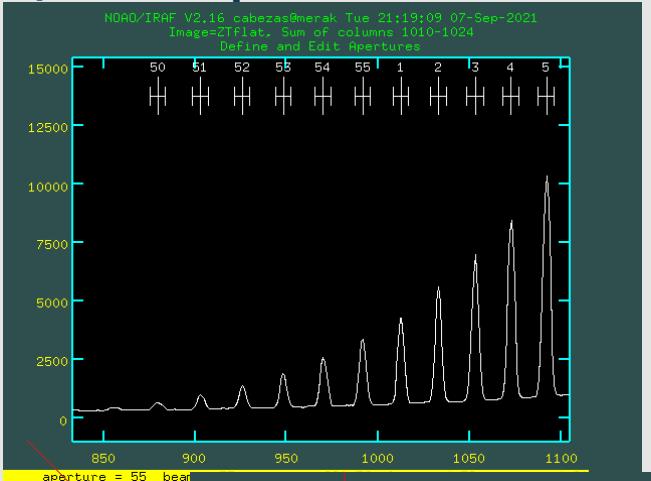
More
Find apertures for ZTflat? (yes):
Number of apertures to be found automatically (49):
Resize apertures for ZTflat? (yes):
Edit apertures for ZTflat? (yes): █
```

8) Model apertures - Flat

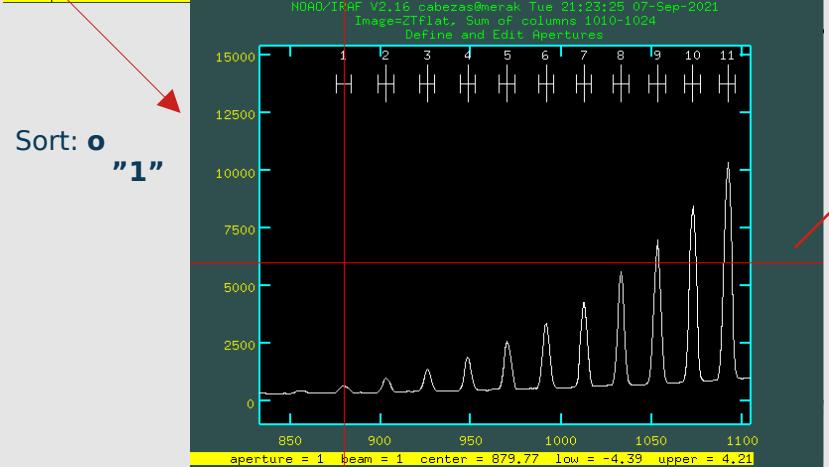
- first aperture near the pixel 890
mark: **m**



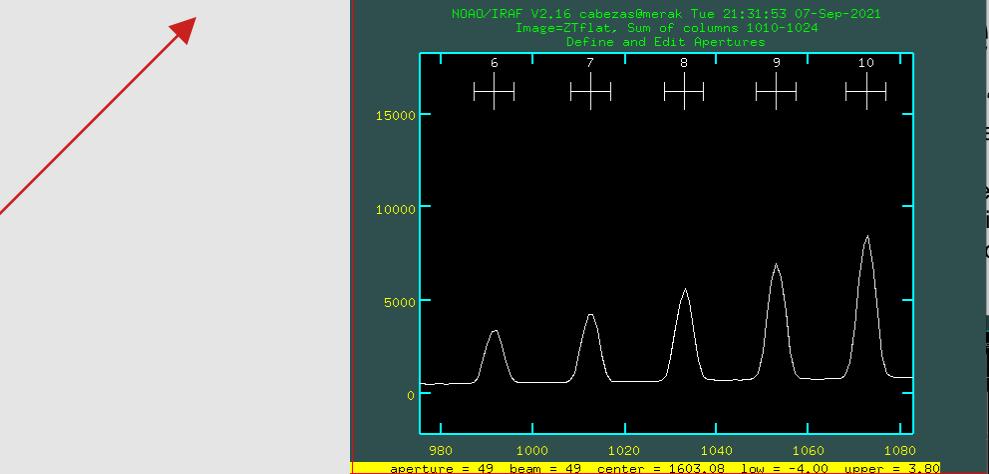
8) Model apertures - Flat



- delete last aperture: **d**



Sort: **o**
"1"



8) Model apertures - Flat

Fit curve to aperture N of Ztflat interactively (yes): YES

Write apertures for ZTflat to database (yes):

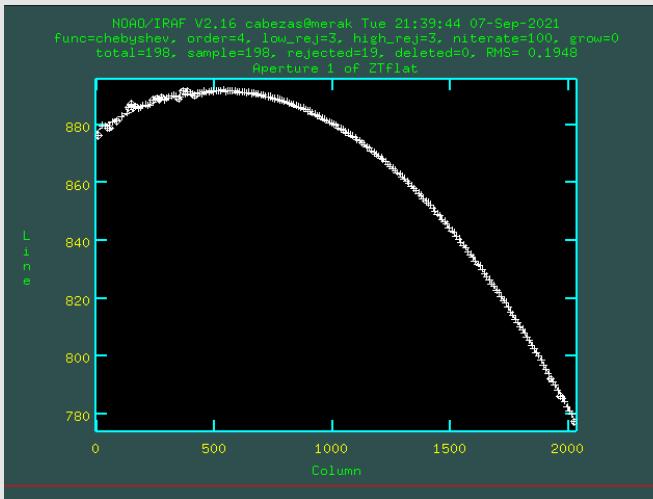
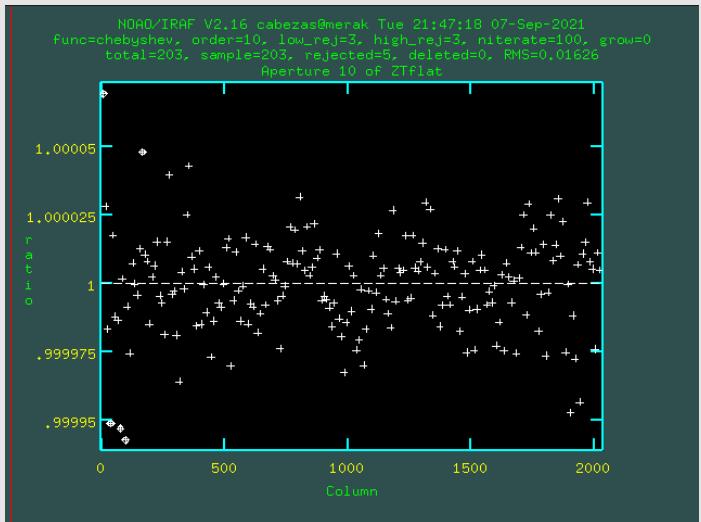
Extract aperture for ZTflat? (yes):

Review extracted spectra from ZTflat? (yes):

Review extracted spectrum for aperture 1 from ZTflat? (yes)

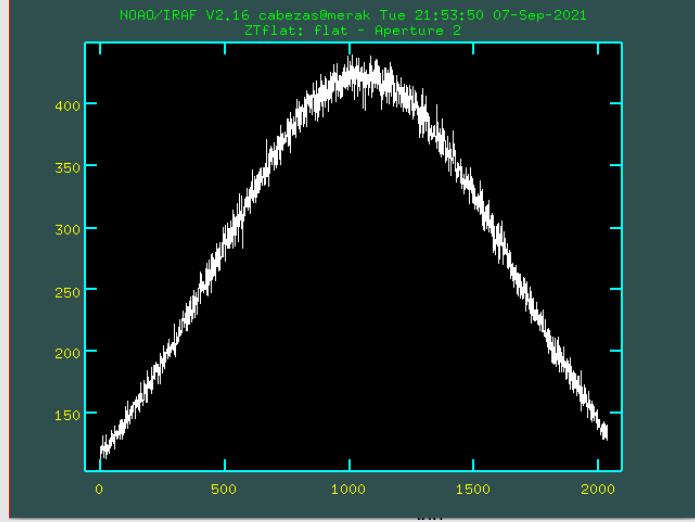
YES

- new file: **AZTflat.fit**



Polynomial fitting of the echelle orders

- change order : **o N**
- relative residuals **k**
- residuals (px) **j**
- aperture **h**



9) Model apertures - Comparison lamp

- Template: AZTflat.fit

iftrimc = yes

flatapa = no

compapa = yes

:go

Edit apertures for Ztflat? (yes):

Can be “no” but always is better to check it!

q

Write apertures for Ztcomp to database (yes):

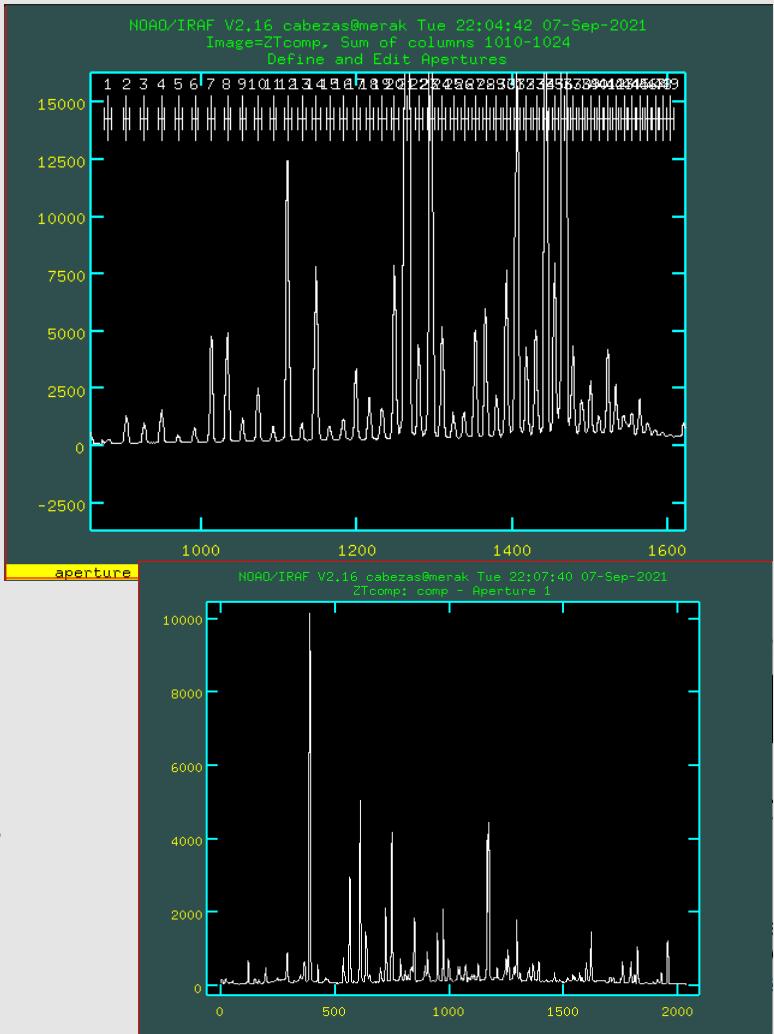
Extract aperture spectra for Ztcomp? (yes):

Review extracted spectra from Ztcomp? (yes):

Review extracted spectrum for aperture 1 from Ztcomp? (yes) YES

- if you are using database, check if the first aperture corresponds to the first aperture in the atlas. (inside folder ***idcomp***)

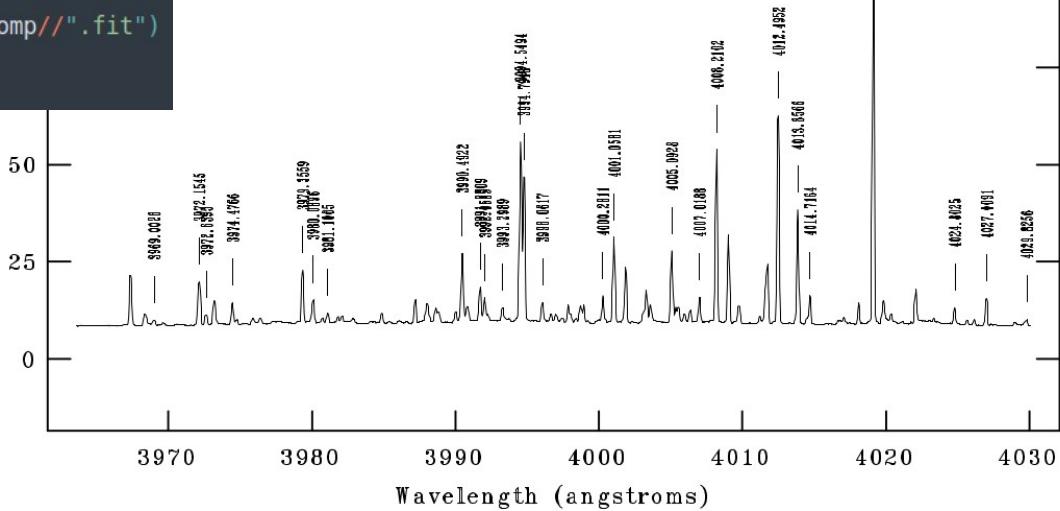
- new file: **AZTcomp.fit**



9) Model apertures - Comparison lamp

```
272 #####APERTURES - COMP
273 if (compapall==yes){
274     apall.referen=inflat
275     apall.format = "echelle"
276     apall.find=no
277     apall.recente=no
278     apall.resize=no
279     apall.trace=no
280     apall.fittrace=no
281     apall.extras=no
282     apall.ylevel = 0.04
283     apall.extract=yes
284     apall (input=incomp, output="A"/incomp//".fit")
285 }
286
```

0/IRAF V2.16 mauricio@mauricio Wed 12:32:54 30-Dec-2020
identify iazcomp.0001 - Ap 1
comp



10) Wavelength calibration

- using database, folder “idcomp”
must be in the main path

iftrimc = yes

compapa = no

iddatab = yes

idfolder = idcomp

idencom = yes

:go

- database can be shifted few pixels,
in order to fix we need to “shift”
the points of our database.

s

- always fit!

f (delete point d), **q, q**

- rms ~ 0.005 acceptable

- ***Write feature data to the
database (yes)?***

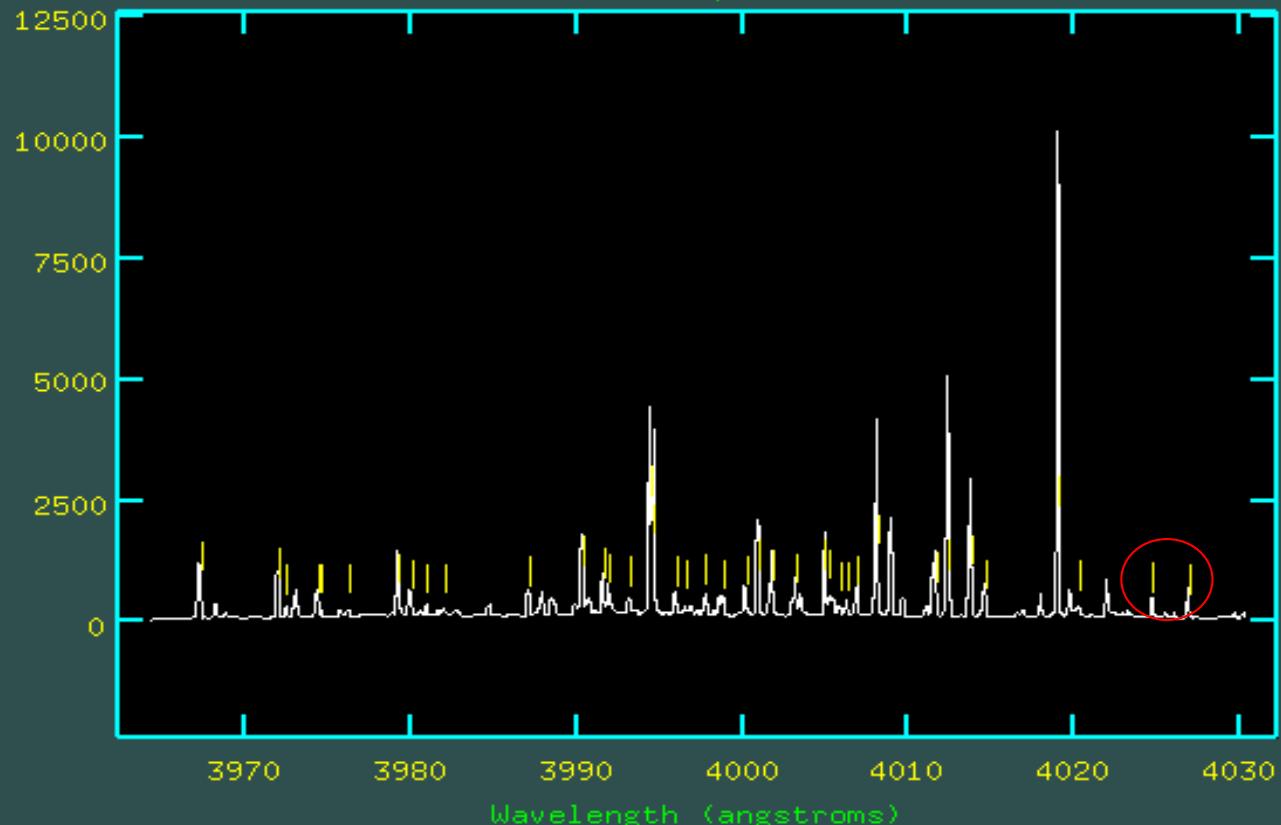
```
I R A F
Image Reduction and Analysis Facility

PACKAGE = clpackage
  TASK = oesred
More
(iddatab=          yes) Use database folder for identification?
(idfolde=          idcomp) folder name with identification database
(idencom=          no) Identify features in spectrum for dispersion sol

# OBJECT PARAMETERS

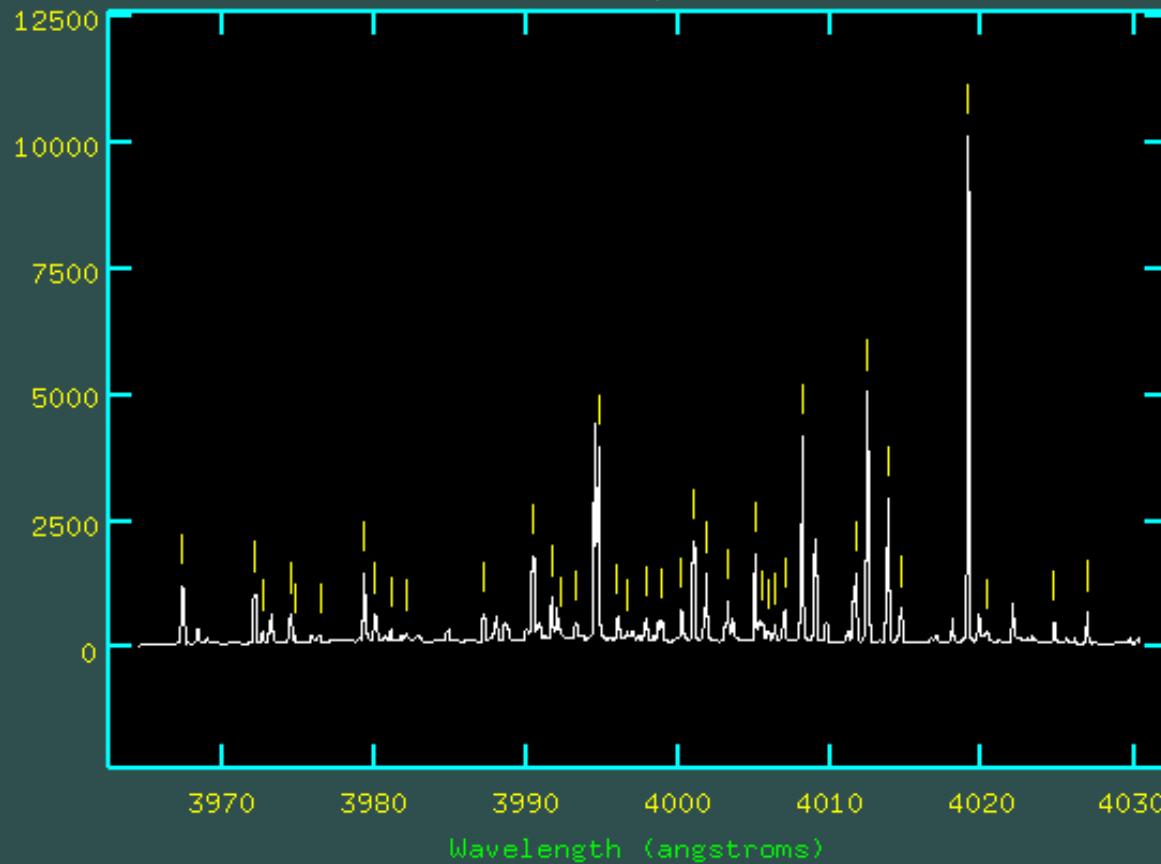
(trimob =          no) Trim object?
(iftrimo=          no) Use trim object?
(zerocor=          no) Apply zero level correction to object?
(crays =           no) Remove cosmic rays?
(ifcrays=          no) Use object with cosmic rays extraction?
(objecta=          no) Extract object apertures?
(flatcor=          no) Apply flat correction to object?
(helioco=          no) calculate JD + heliocentric correction?
(idref =           no) refer database identification to images?
(norm   =           no) normalize spectra?
(ncombin=          no) combine normalized spcectra?
More
ESC-? for HELP
```

NOAO/IRAF V2.16 cabezas@merak Tue 22:21:55 07-Sep-2021
identify iazcomp.0001 - Ap 1
comp



No match!
Press "s"

NOAO/IRAF V2.16 cabezas@merak Tue 22:24:31 07-Sep-2021
identify iazcomp.0001 - Ap 1
comp



Recentered=39/39, pixel shift=3.33, user shift=-0.11, z=-2.7E-5, rms=0.057

11) Trim Object

iftrimc = yes

iddatab = no

idfolder = idcomp

idencom = no

trimob = yes

iftrimo = yes

:go

New file: **Te202109060016.fit**

```
I R A F
Image Reduction and Analysis Facility
PACKAGE = clpackage
TASK = oesred
More
(iddatab=           no) Use database folder for identification?
(idfolde=          idcomp) folder name with identification database
(idencom=           no) Identify features in spectrum for dispersion sol

# OBJECT PARAMETERS

(trimob=           yes) Trim object?
(iftrimo=          yes) Use trim object?
(zerocor=           no) Apply zero level correction to object?
(crays=             no) Remove cosmic rays?
(ifcrays=           no) Use object with cosmic rays extraction?
(objecta=           no) Extract object apertures?
(flatcor=           no) Apply flat correction to object?
(helioco=           no) calculate JD + heliocentric correction?
(idref=              no) refer database identification to images?
(norm=               no) normalize spectra?
(ncombin=            no) combine normalized spcectra?
More
ecl> []
```

The screenshot shows a terminal window with the IRAF command-line interface. The command history at the bottom includes:

- as on...asu.cas.cz
- 20210906
- alpLyr ▾
- Q
- ≡
-
- ×

The file listing below shows three files:

- e20210906 0016.fit
- logfile
- Te20210906 0016.fit

12) Bias correction object

iftrimc = yes

trimob = no

iftrimo = yes

zerocor = yes

:go

New file: **ZTe202109060016.fit**

```
I R A F
Image Reduction and Analysis Facility

PACKAGE = clpackage
  TASK = oesred
More
(iddatabase=          no) Use database folder for identification?
(idfolde=            idcomp) folder name with identification database
(idencom=            no) Identify features in spectrum for dispersion sol

# OBJECT PARAMETERS

(trimob =           no) Trim object?
(iftrimo=          yes) Use trim object?
(zerocor=          yes) Apply zero level correction to object?
(cray =             no) Remove cosmic rays?
(ifcray=            no) Use object with cosmic rays extraction?
(objecta=           no) Extract object apertures?
(flatcor=           no) Apply flat correction to object?
(helioco=            no) calculate JD + heliocentric correction?
(idref =             no) refer database identification to images?
(norm =              no) normalize spectra?
(ncombin=            no) combine normalized spcectra?
More
ecl> □

  s on...asu.cas.cz  20210906  alpLyr ▾
  Q  E  ▾  ▾  -  □  X

  e202109060016.fit  logfile  Te202109060016.fit  ZTe202109060016.fit
```

13) Cosmic Rays - LACOS (2001PASP.113.1420V)

iftrimc = yes
iftrimo = yes
zerocor = no
crays = yes
ifcrays = yes
:go

New file: **CrZTe202109060016.fit**

Some issues with parameters

Best setup?

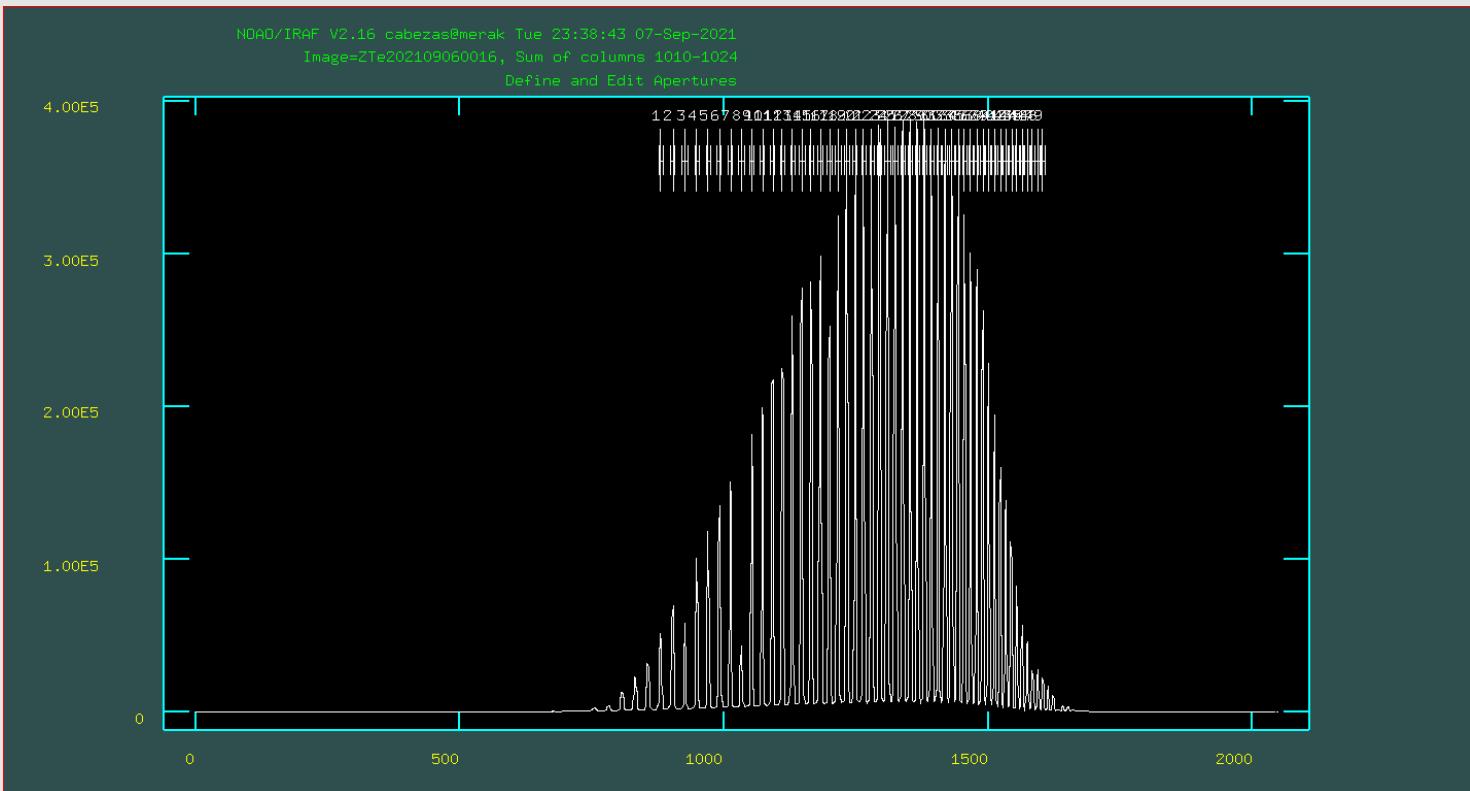
Different for each target/night/exposure
<http://www.astro.yale.edu/dokkum/lacosmic/>

```
400 #####COSMIC RAYS - COMP
401 if (crays==yes){
402     stsdas
403 #####
404 #read gain
405 cd (oname)
406 hselect (images="Z"//inobject,fields="GAIN", exp=yes) | scan (gainh)
407 hselect (images="Z"//inobject,fields="READNOIS", exp=yes) | scan (readnh)
408 #print (gainh)
409 inputCR="Z"//inobject
410 outputCR="CrZ"//inobject
411 outmaskCR="MCrZ"//inobject
412 gainCR = gainh # 2 #3
413 readnCR = readnh #2
414 xorderCR = 3
415 yorderCR = 3
416 sigclipCR = 4.5
417 sigfracCR = 0.3
418 objlimCR = 4
419 niterCR = 5
420 verboseCR = no
421 ..
```

14) Model apertures - Object

- Template: AZTflat.fit

```
iftrimc = yes  
iftrimo = yes  
crays = no  
ifcrays = no  
objecta = yes  
:go
```



14) Model apertures - Object

Edit apertures for ZTe202109060016? (yes):

Can be “no” but always is better to check it!

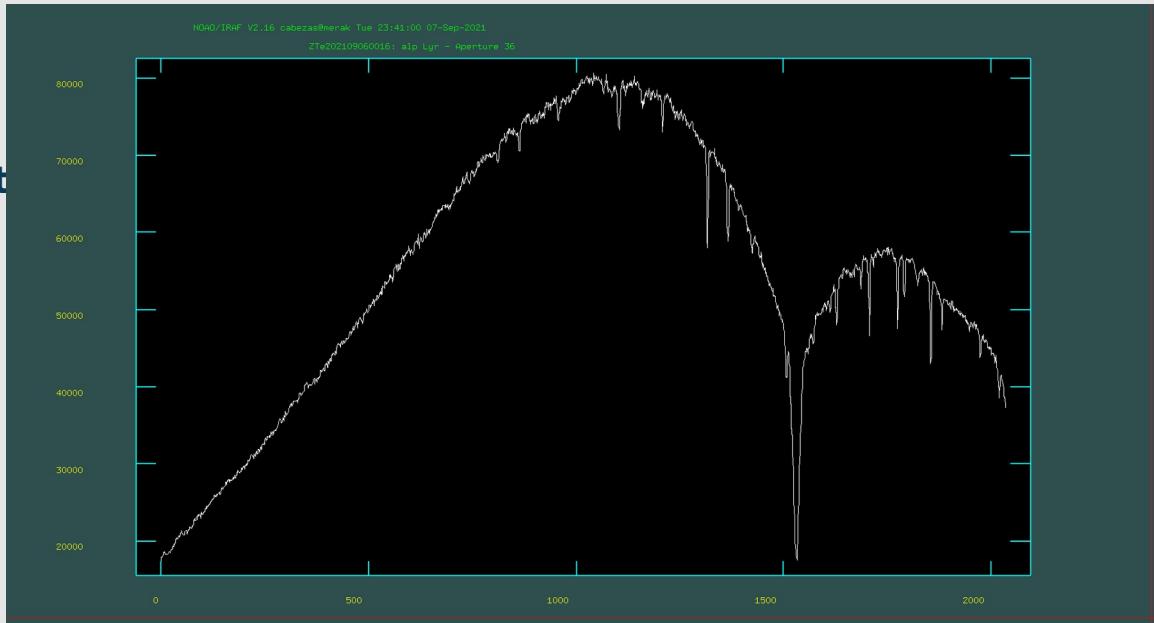
q

Review extracted spectrum for aperture 1 from

ZTe202109060016? (yes) YES

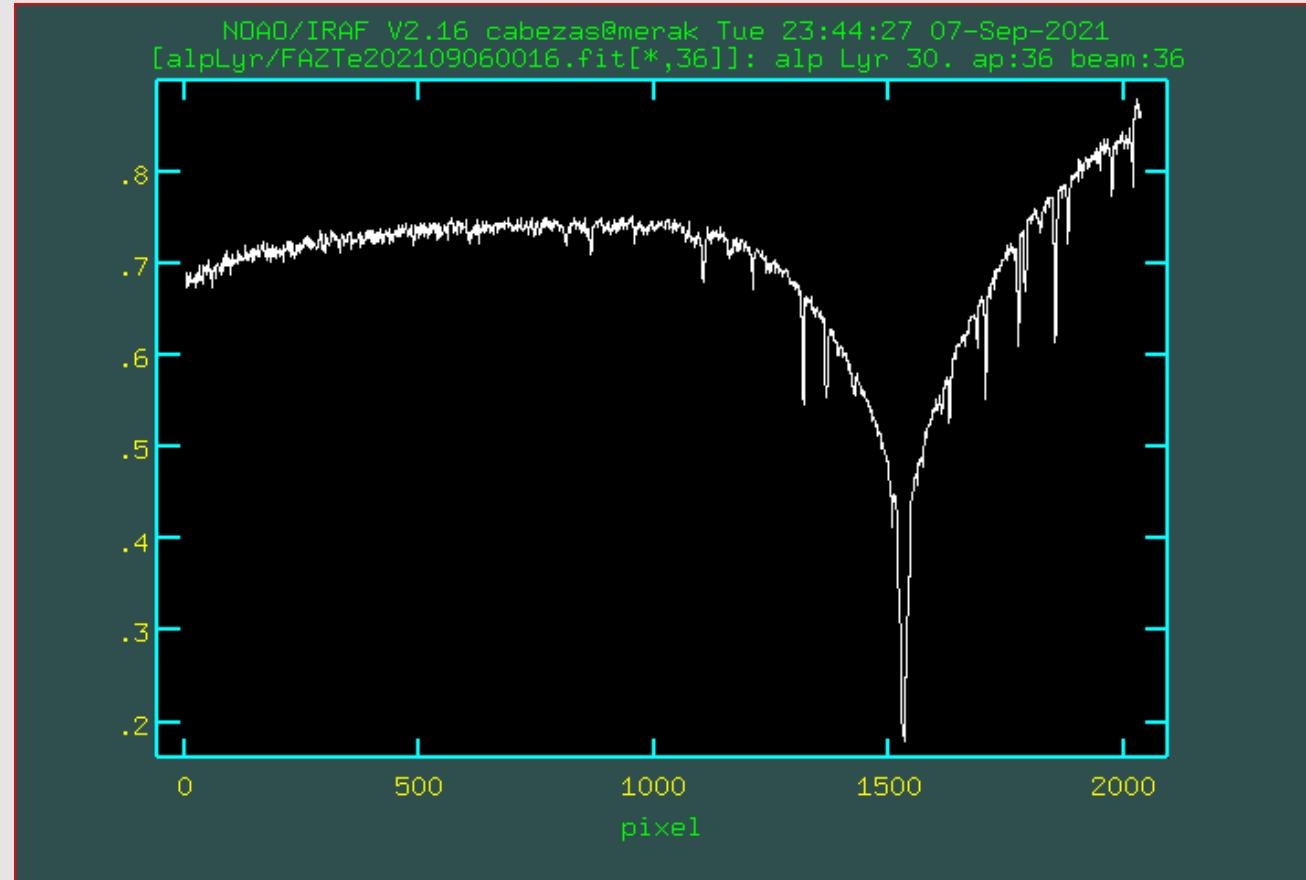
- still each spectrum in pixels

new file: ***AZTe202109060016.fit***



15) Flat correction

```
iftrimc = yes  
iftrimo = yes  
ifcrays = no  
objecta = no  
flatcor = yes  
:go
```



new file: **FAZTe202109060016.fit**

16) JD + heliocentric correction

```
iftrimc = yes          (crays =           no) Remove cosmic rays?
iftrimo = yes          (ifcrays=         no) Use object with cosmic rays extraction?
ifcrays = no           (objecta=        no) Extract object apertures?
flatcor = no           (flatcor=        no) Apply flat correction to object?
helioco = yes          (helioco=       yes) calculate JD + heliocentric correction?
idref   =               (idref =         no) refer database identification to images?
norm    =               (norm =         no) normalize spectra?
ncombin=              (ncombin=      no) combine normalized spcectra?
More
18:9:0ZTe202109060016.fit
#             Image          jd          hjd          ljd
# SETJD: Observatory parameters for Ondrejov observatory
#     timezone = -1
Warning: Image header parameter not found (UTMIDDLE)
# RVCORRECT: Observatory parameters for Ondrejov observatory
#     latitude = 49:54:38
#     longitude = 345:12:59
#     altitude = 528
##YR MO DY    UT      RA      DEC      VOBS
## HJD          VOBS    VHELIO    VLSR    VDIURNAL   VLUNAR   VANNUAL   VSOLAR
2021  9  6 18:09:00 18:36:56 38:47:01    0.0
2459464.25768    0.00  -12.05     7.57     0.026   -0.005  -12.072  19.625
FAZTe202109060016.fit
imhead FAZTe202109060016.fit I+ 2459464.25768 -12.05
```

```
CDZ_Z = 
HJD = 2459464.25767855
VHELIO = -12.0510150041803
VLSR = 7.57409507948893
VSUN = ' 20.    18.    30.  1900.'
ecl>
```

16) JD + heliocentric correction

New UTMIDDLE

```
347 #CALC UTMIDDLE
348 hselect (images="other://"//spec,fields="TM_START", exp=yes) | scan (utstart)
349 hselect (images="other://"//spec,fields="EXPTIME", exp=yes) | scan (expt)
350 utmidhr=int((utstart + expt/2)/3600)
351 utmidmin=int(((utstart + expt/2)/3600)-utmidhr)*60
352 utmidsec=int((((utstart + expt/2)/3600)-utmidhr)*60 - utmidmin)*60
353 utmid = (utmidhr//":"//utmidmin//":"//utmidsec)
354 printf ("%d:%d:%d\n",utmidhr,utmidmin,utmidsec) | scan (utmid)
355 #####
```

17) Ref spectrum

iftrimc = yes

iftrimo = yes

ifcrays = no

helioco = no

idref = yes

:go

```
(ncombin=                               no) combine normalized spcectra?
More
[ap.0001] refspec1='iazcomp.0001'
ap.0001.fit: REFSPEC1 = 'iazcomp.0001 1.'
wap.0001.fit: ap = 1, w1 = 3964.476, w2 = 4030.398, dw = 0.032426, nw = 2034
[ap.0002] refspec1='iazcomp.0002'
ap.0002.fit: REFSPEC1 = 'iazcomp.0002 1.'
wap.0002.fit: ap = 2, w1 = 4009.502, w2 = 4076.196, dw = 0.032806, nw = 2034
[ap.0003] refspec1='iazcomp.0003'
ap.0003.fit: REFSPEC1 = 'iazcomp.0003 1.'
wap.0003.fit: ap = 3, w1 = 4055.563, w2 = 4123.023, dw = 0.033183, nw = 2034
[ap.0004] refspec1='iazcomp.0004'
ap.0004.fit: REFSPEC1 = 'iazcomp.0004 1.'
wap.0004.fit: ap = 4, w1 = 4102.691, w2 = 4170.963, dw = 0.033582, nw = 2034
[ap.0005] refspec1='iazcomp.0005'
ap.0005.fit: REFSPEC1 = 'iazcomp.0005 1.'
wap.0005.fit: ap = 5, w1 = 4150.934, w2 = 4220.016, dw = 0.03398, nw = 2034
[ap.0006] refspec1='iazcomp.0006'
ap.0006.fit: REFSPEC1 = 'iazcomp.0006 1.'
wap.0006.fit: ap = 6, w1 = 4200.32, w2 = 4270.232, dw = 0.034389, nw = 2034
[ap.0007] refspec1='iazcomp.0007'
ap.0007.fit: REFSPEC1 = 'iazcomp.0007 1.'
wap.0007.fit: ap = 7, w1 = 4250.877, w2 = 4321.653, dw = 0.034813, nw = 2034
[ap.0008] refspec1='iazcomp.0008'
```

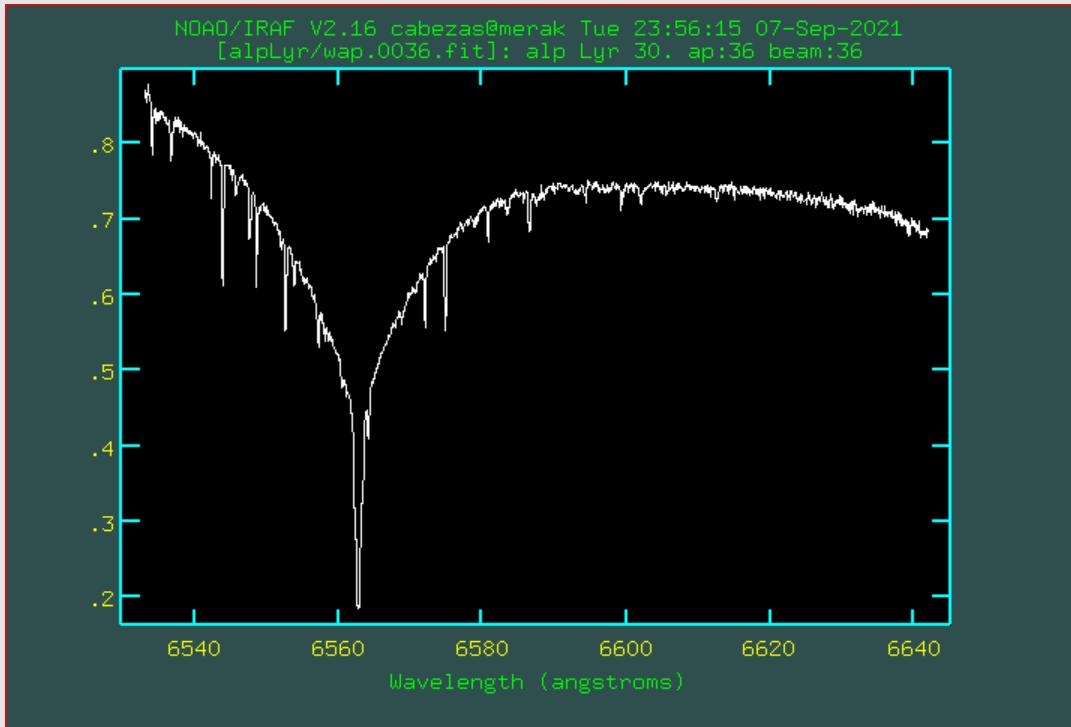
17) Ref spectrum

Read database of wavelength calibration and apply!

New files:

ap.00XX.fit - 1D spectra in pixels

wap.00XX.fit - 1D spectra in Angstrom



18) Normalization

```
iftrimc = yes  
iftrimo = yes  
ifcrays = no  
idref = no  
norm = yes  
:go
```

Fit [1,1] of wap.0001.fit w/ graph?
(yes/no/skip/YES/NO/SKIP) (yes):

```
# OBJECT PARAMETERS

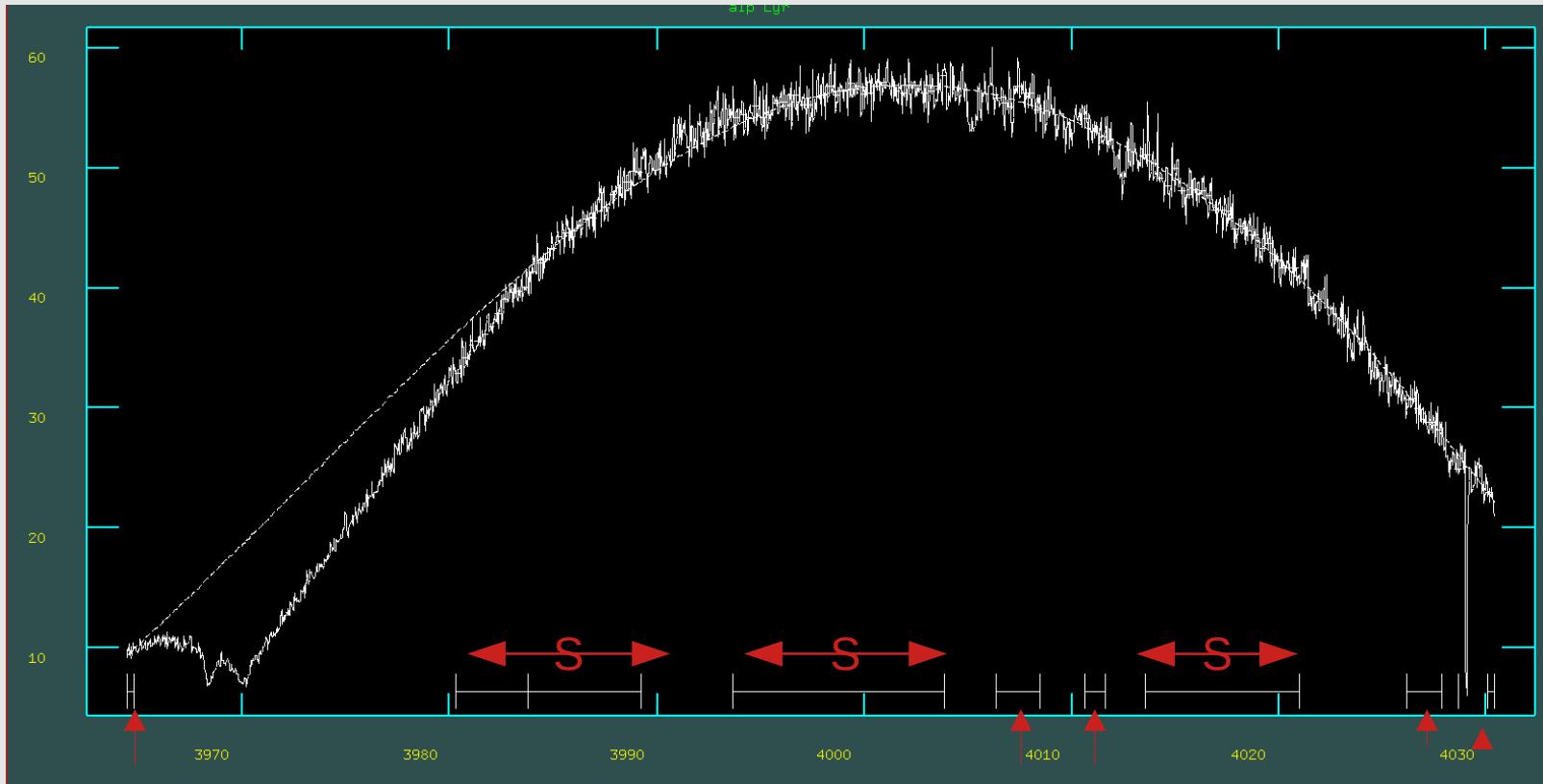
(trimob =
(iftrimo=
(zerocon=
(crays =
(ifcrays=
(objecta=
(flatcor=
(helioco=
(idref =
(norm =
(ncombin=
More
wap.0001.fitfap.0001.fitnap.0001.fit
Fit [1,1] of wap.0001.fit w/ graph? (yes|no|skip|YES|NO|SKIP) (yes):
```

18) Normalization

- change order :**o N**
- residuals **j**
- high rejection :**hi N**

- low rejection :**lo N**
- function :**f legendre/spline3/chebyshev**
- select region **s..s**
- delete region **z**
- delete all regions **t**

- zoom/window **w, e..e**
- resize **w,a**
- move right **w,r**
- move left **w,l**
- move up **w,u**
- move down **w,d**



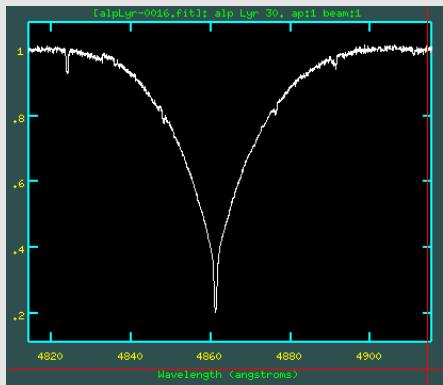
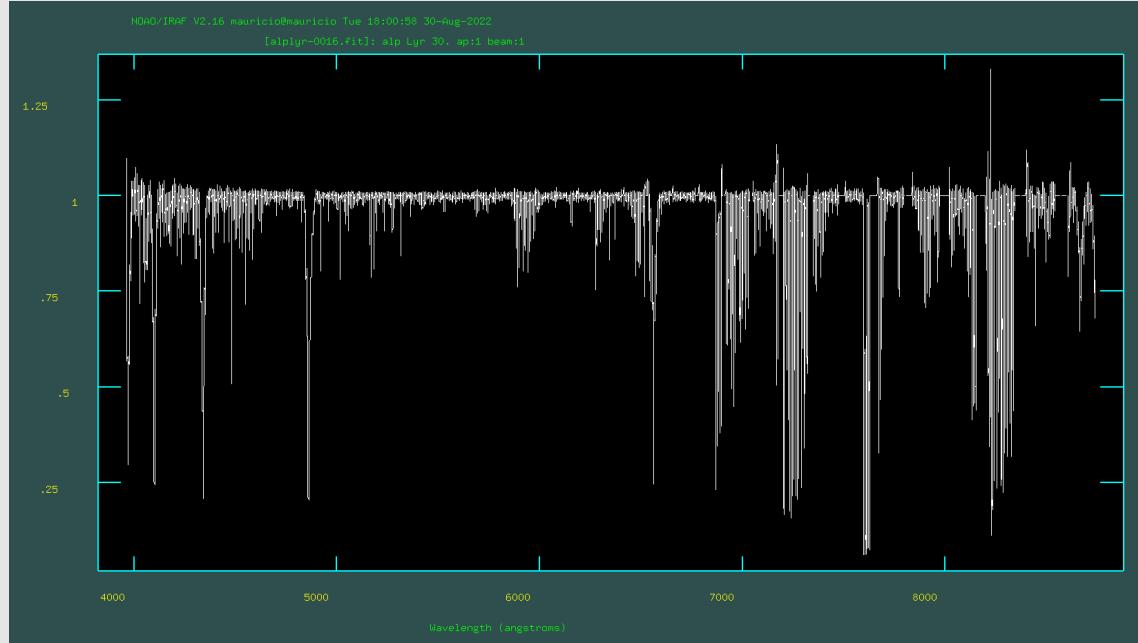
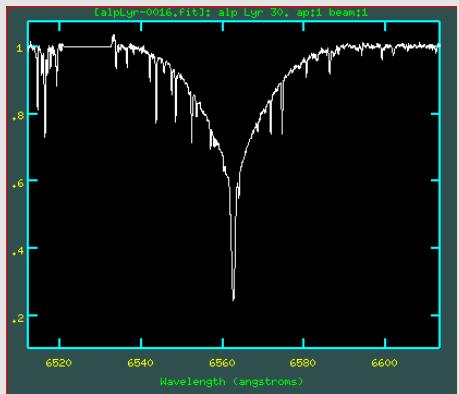
18) Normalization

```
839
840 #norm=no
841 ##### normalization
842 if (norm==yes){
843   cd (oname)
844   unlearn continuum
845   unlearn scombine
846   continuum.type="fit"
847   continuum.function=cfunction
848   continuum.order=corder
849   continuum.naverage=10
850   continuum.markrej=no
851   continuum.niterat=2000
852   continuum.high_re=2
853   continuum.low_re=1.5
854   continuum.grow=0
855   for (i=1; i <=nap; i+=1) {
856     printf ("wap.00%02d.fit\n", (i)) | scan(wap)
857     printf ("fap.00%02d.fit\n", (i)) | scan(fap)
858     printf ("nap.00%02d.fit\n", (i)) | scan(nnap)
859     print (wap, fap, nnap)
860   continuum (input=wap, output=fap)
861 }
862   cd "../"
863 }
```

21) merging

```
iftrimc = yes  
iftrimo = yes  
ifcrays = no  
norm = no  
ncombine = yes  
:go
```

new file: alpLyr-0016.fit



S P L O T

splot alpLyr-0016.fit

Fit: gaussian: **k..k(or g)**

lorentzian: **k..l**

voigt: **k..v**

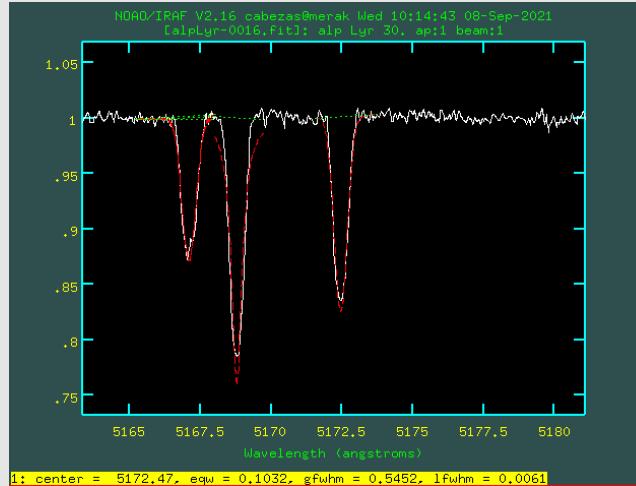
centroid **e..e**

snr: **m..m**

Change unit (angstrom to km/s)

:u km/s 6562.8 an

:u an



S P E C P L O T

specplot spec1.fit,spec2.fits

Change step: **step 1** (or any number)

See wavelength: **u**

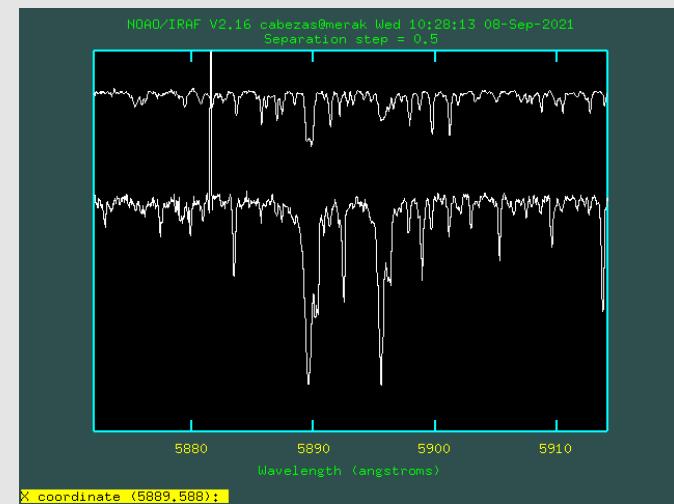
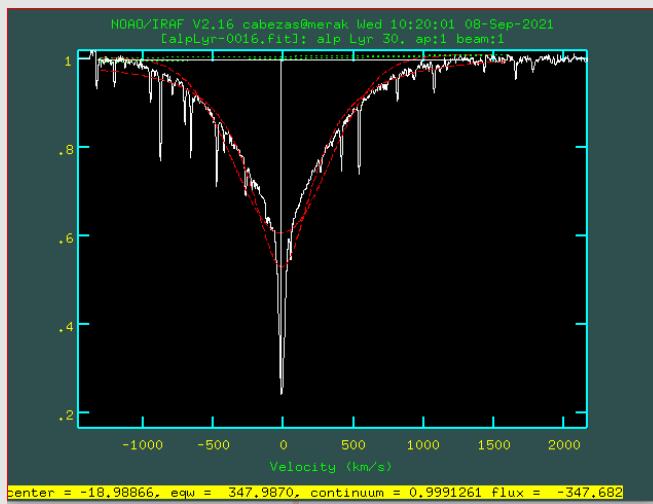
snr: **m..m**

Change unit (angstrom to km/s)

:u km/s 6562.8 an

:u an

Replot: **r**



General Remarks

Thank
you!!

- The optimal reduction process always will be different for each instrument.
- IRAF “sometimes” is a bit tricky, but really useful.
- Quick check/inspection of spectra!!
- Versatil program because many parameters (sometimes too much).
- Pre-defined task.
- “opensource” you can write your own task/package.
- Xgterm - nice interactive tool.
- Slow with computation, python/idl/fortran would be good option.