

Astronomical Virtual Observatory and its Scientific Challenges

Petr Škoda

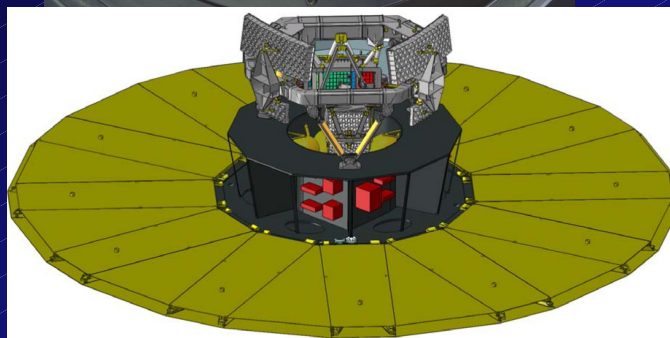
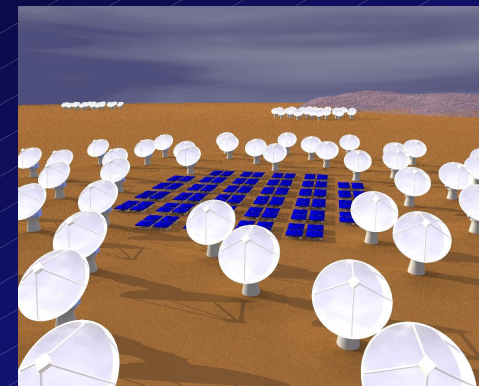
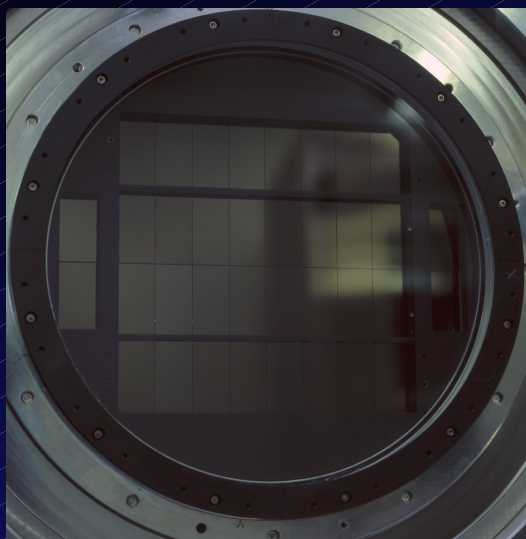
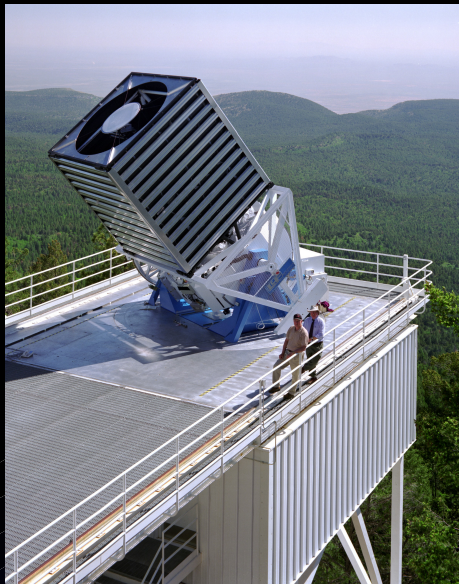
Astronomical Institute Academy of Sciences
Ondřejov
Czech Republic

Institute of Astronomy of BAS, Sofia, Bulgaria, 1-st December 2009

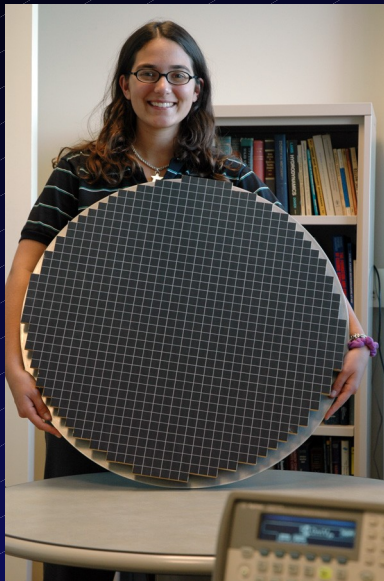
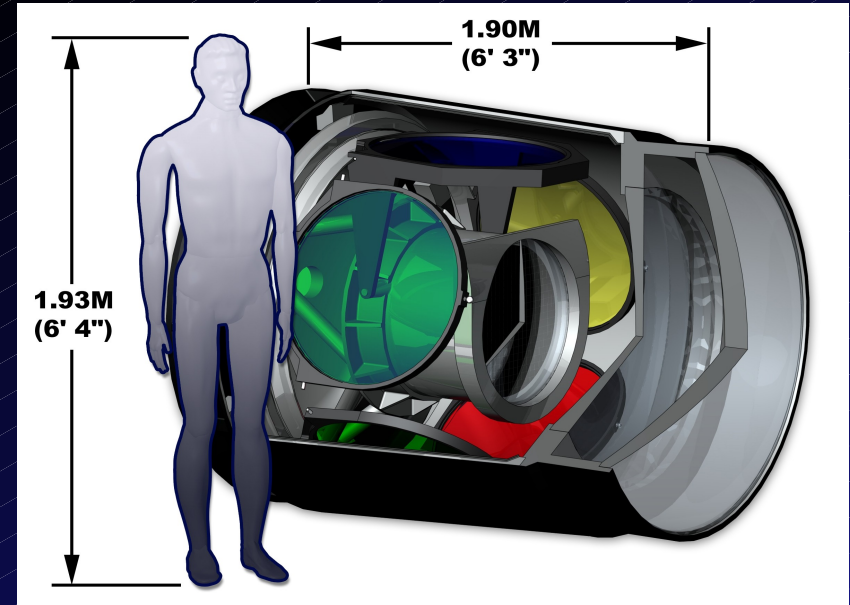
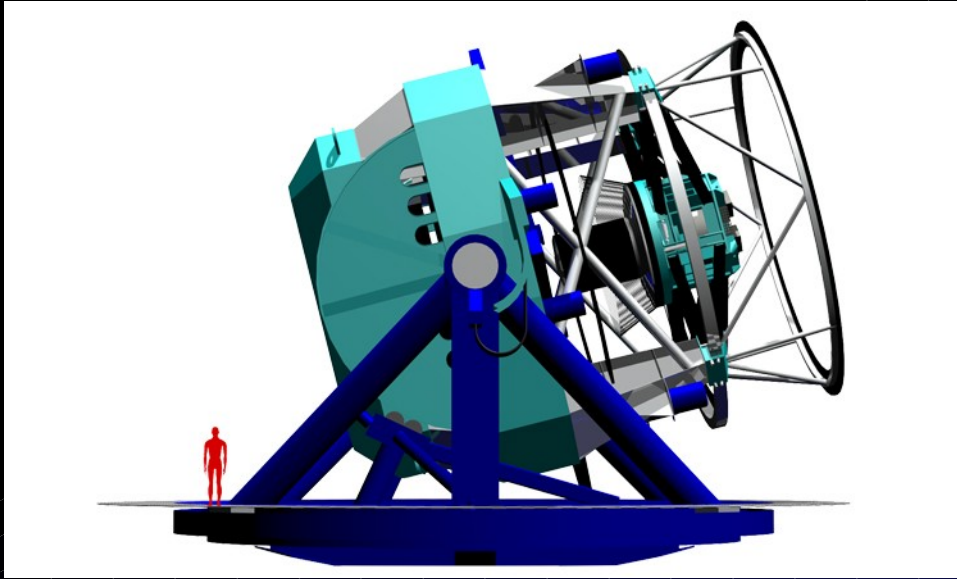
Data Avalanche

- Huge surveys: 100 million sources at < 3000 sources per night $\Rightarrow$ > 100 years to identify them
- Huge data collections: download and data analysis on desktop problematic/impossible.
- Example: downloading Sloan Digital Sky Survey (SDSS) DR6 data:
 - ✓ images (10 Terabytes) $\Rightarrow$ ~ 3 months at 10 Mbps
 - ✓ catalogues (2 Terabytes) $\Rightarrow$ ~ 3 weeks
 - ✓ on DVDs $\Rightarrow$ ~ 2,100 of them
- And data analysis?? (similar size for MACHO, 2MASS et

Data Avalanche



LSST (8.4m)

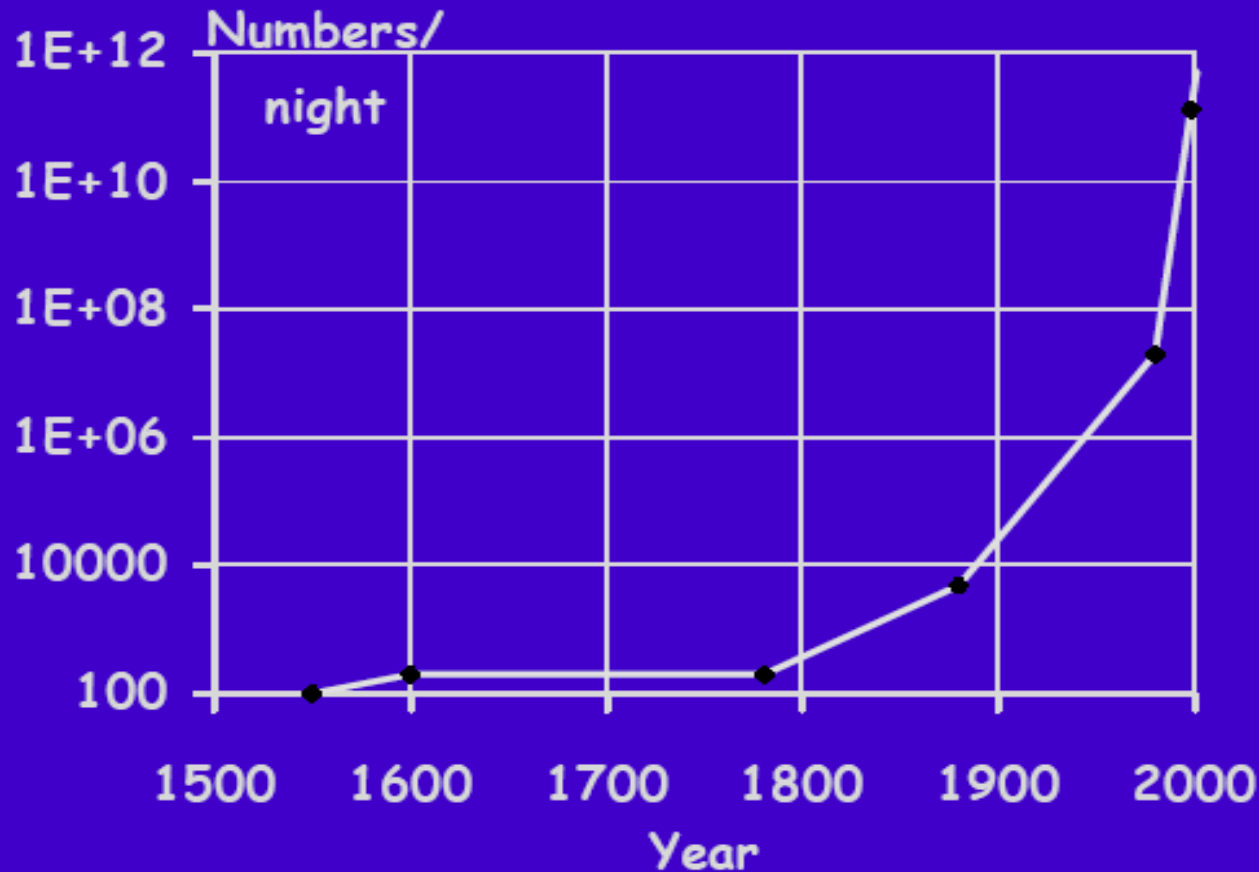


200 CCD 4kx4k,
32 channels (6400)
3.2 Gpix every 20 sec
64cm diameter
3.5 deg FOV
30 TB/night
2 TFLOPS
detection of changes
within 60sec

Data Avalanche

Moore law for chips –doubling 1.5 year (1000/10 yr)

Data in astronomy – doubling < 1 yr !



$T_2 < 18$ mths
1990-2000

History of VO

idea VO end 2000

Federation of archives (MAST)

unified IF, data format for transport

much data – distributed processing

GRID - started in HEP

Multispectral research : radio---gama

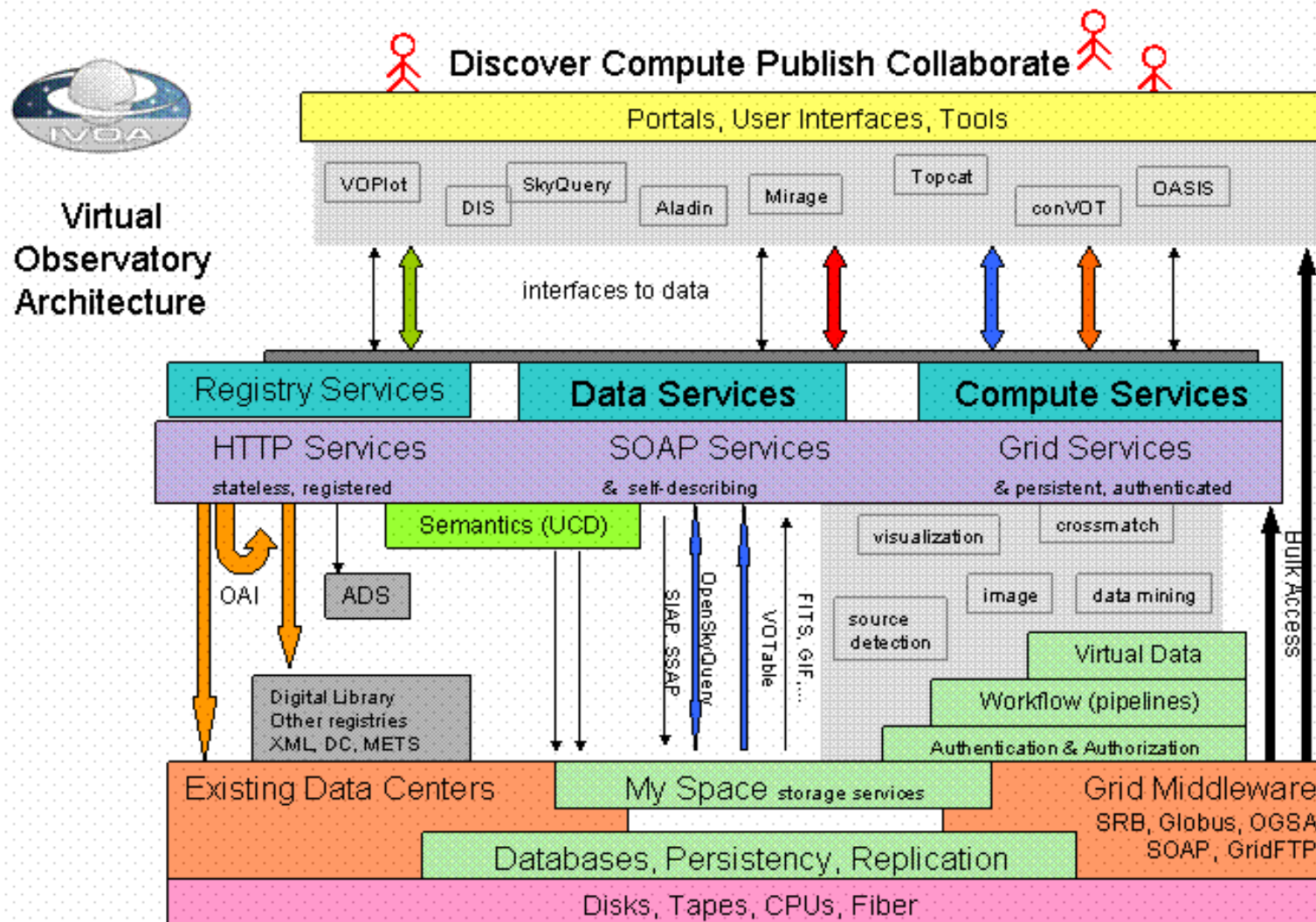
Output of simulations in NCSA

Data for SDSS – key research

VO Paradigm

- METADATA (name of column), ontologies (name)
- Unique format (VOTable – e.g VizieR)
- Transparent search, download, conversion
- Query for data – processing done on servers
- Federation of astronomical archives (protocols)
- Unified presentation – automatic units conversion (A, MeV, MHz $\rightarrow$ nm), $\text{Wm}^{-2}\text{s}^{-1} \rightarrow \text{Jy}$)
- Background computing on GRIDS
- Multiwavelength approach (SED)

Architecture of VO



Technology of VO

Unified data format– VOTable, UCD

Transparent transport (SOAP , REST)

Web services (WS)

VOregistry (like DNS)

VOspace (network home disc)

protocols (CGI services)

ConeSearch (searching in circle on sky)

SIAP (Simple Image Access Protocol)

SSAP(Simple Spectral Access Protocol)

SLAP(Simple Line Access Protocol)

TAP (Table Access Protocol)

Technology of VO

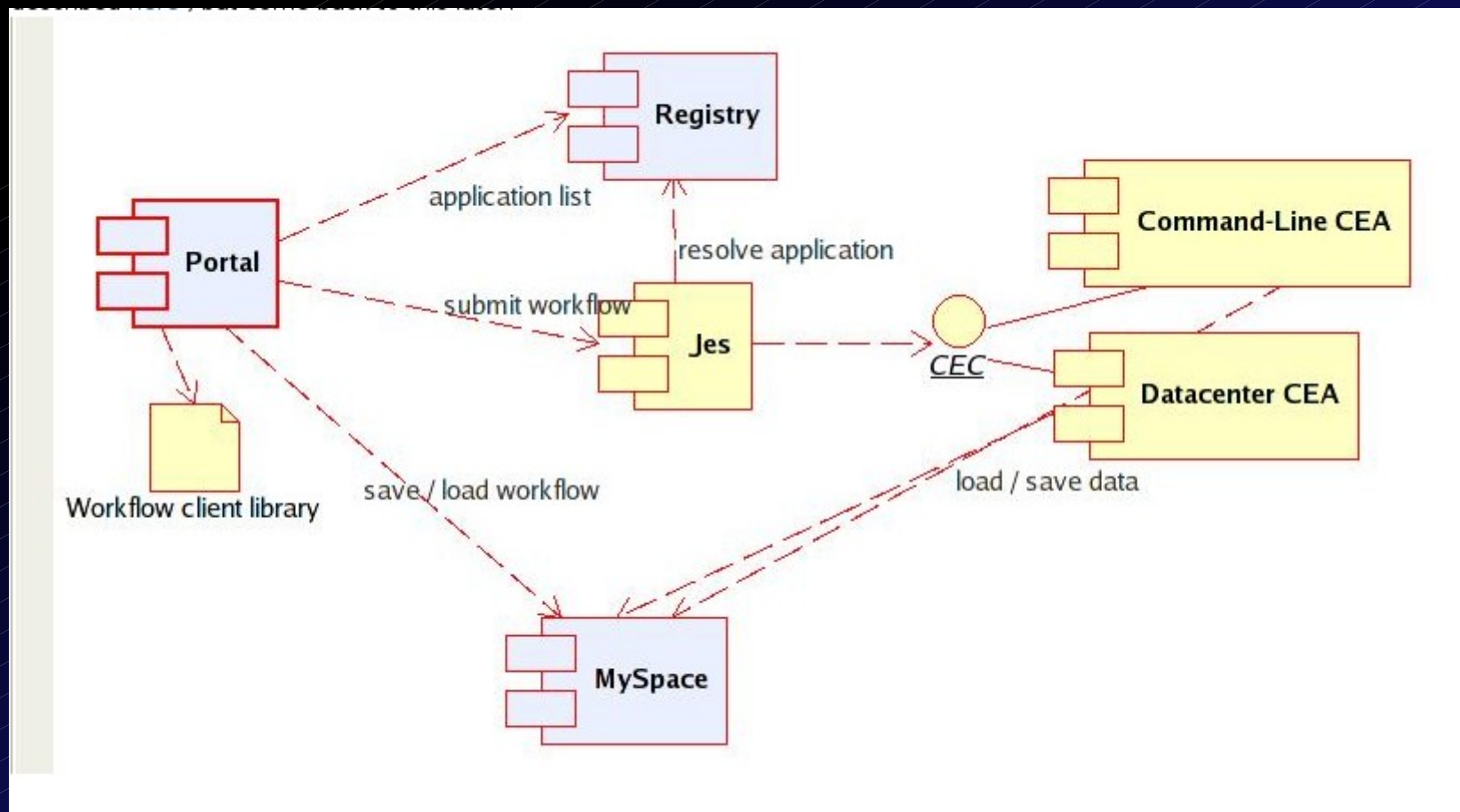
ADQL

XMATCH, REGION

VOSpace – cache(WS, database, http put get)

Application interop – PLASTIC, SAMP

Workflows - Astrogrid



**VOTable :**

```
<?xml version="1.0"?>
<!DOCTYPE VOTABLE SYSTEM "http://us-vo.org/xml/VOTable.dtd">
<VOTABLE >
  <DESCRIPTION>
    VizieR Astronomical Server: urania.iucaa.ernet.in 2002-10-04T05:20:16
    Explanations and Statistics of UCDs: See LINK below
    In case of problem, please report to: question@simbad.u-strasbg.fr
  </DESCRIPTION>
  <DEFINITIONS>
    <COOSYS ID="J2000" equinox="J2000" system="EQ_FK5"/>
  </DEFINITIONS>
  <INFO ID="Ref" name="-ref" value="VOTx11451"/>
  <RESOURCE name="V/105" ID="yCat_5105" >
    <DESCRIPTION>SKY2000 Catalog, Version 3 (Myers+ 2000)
  </DESCRIPTION>
  <TABLE ID="V_105_sky2v3r1" name="V/105/sky2v3r1" >
    <DESCRIPTION>The Sky2000 Version 2 Catalogue
  </DESCRIPTION>
    <FIELD datatype="int" width="6" name="HD" ucd="ID_ALTERNATIVE" >
      <DESCRIPTION>?Henry Draper &lt;III/135> number
    </DESCRIPTION>
    </FIELD>
    <FIELD unit="h:m:s" datatype="char" ref="J2000" name="RAJ2000" ucd="POS_EQ_RA_MAIN" arraysize="13" >
      <DESCRIPTION>Right ascension (J2000) hours
    </DESCRIPTION>
    </FIELD>
    <FIELD unit="d:m:s" datatype="char" ref="J2000" name="DEJ2000" ucd="POS_EQ_DEC_MAIN" arraysize="13" >
      <DESCRIPTION>Declination degrees (J2000)
    </DESCRIPTION>
```

☒ Display Data Of Selected Points

Close

Save As File

Simple Spectra Access Protocol Spectral Data Model

Simple Spectral Access Protocol V1.04



*International
Virtual
Observatory
Alliance*

Simple Spectral Access Protocol

Version 1.04

IVOA Recommendation Feb 01, 2008

This version:

<http://www.ivoa.net/Documents/REC/DAL/SSA-20080201.html>

Latest version:

<http://www.ivoa.net/Documents/latest/SSA.html>

Previous version(s):

Version 1.03, December 2007
Version 1.02, September 2007
Version 1.01, June 2007
Version 1.00, May 2007
Version 0.97, November 2006
Version 0.96, September 2006
Version 0.95 May 2006
Version 0.91 October 2005
Version 0.90 May 2005

Editors:

D.Tody, M. Dolensky

Authors:

D.Tody, M. Dolensky, J. McDowell, F. Bonnarel, T. Budavari, I. Busko, A. Micol, P. Osuna, J. Salgado, P. Skoda, R. Thompson, F. Valdes, and the data access layer working group.



*International
Virtual
Observatory
Alliance*

IVOA Spectral Data Model

Version 1.03

IVOA Recommendation 2007-10-29

This version (Recommendation Rev 1)

<http://www.ivoa.net/Documents/REC/DM/SpectrumDM-20071029.pdf>

Latest version:

<http://www.ivoa.net/Documents/latest/SpectrumDM.html>

Previous versions:

<http://www.ivoa.net/Documents/PR/DM/SpectrumDM-20070913.html>

Editors:

Jonathan McDowell, Doug Tody

Contributors:

Jonathan McDowell, Doug Tody, Tamas Budavari, Markus Dolensky, Inga Kamp, Kelly McCusker, Pavlos Protopapas, Arnold Rots, Randy Thompson, Frank Valdes, Petr Skoda, and the IVOA Data Access Layer and Data Model Working Groups.

SSAP Parameters

4.1.1 Mandatory Query Parameters

The following parameters **must** be implemented by a compliant service:

Parameter	Sample value	Physical unit	Datatype
POS	52, -27.8	degrees; defaults to ICRS	string
SIZE	0.05	degrees	double
BAND	2.7E-7/0.13	meters	string
TIME	1998-05-21/1999	ISO 8601 UTC	string
FORMAT	votable	-	string

4.1.2 Recommended and Optional Query Parameters

Parameter	Sample value	Unit	Req	Datatype
APERTURE	0.00028 (=1")	degrees	OPT	double
SPECRP	2000	$\lambda/d\lambda$	REC	double
SPATRES	0.05	degrees	REC	double
TIMERES	31536000 (=1yr)	seconds	OPT	double
SNR	5.0	dimensionless	OPT	double
REDSHIFT	1.3/3.0	dimensionless	OPT	string
VARAMPL	0.77	dimensionless	OPT	string
TARGETNAME	mars		OPT	string
TARGETCLASS	star		OPT	string
FLUXCALIB	relative		OPT	string
WAVECALIB	absolute		OPT	string
PUBDID	ADS/col#R5983		REC	string
CREATORID	ivo://auth/col#R1234		REC	string
COLLECTION	SDSS-DR5		REC	string
TOP	20	dimensionless	REC	int
MAXREC	5000		REC	string
MTIME	2005-01-01/2006-01-01	ISO 8601	REC	string
COMPRESS	true		REC	boolean
RUNID			REC	string

The spatial, spectral, and time resolution of the data may all be used as query parameters.

IVOA



VO-enabled tools

Aladin

VOPlot

TOPCAT

VOSpec

SpecView

SPLAT

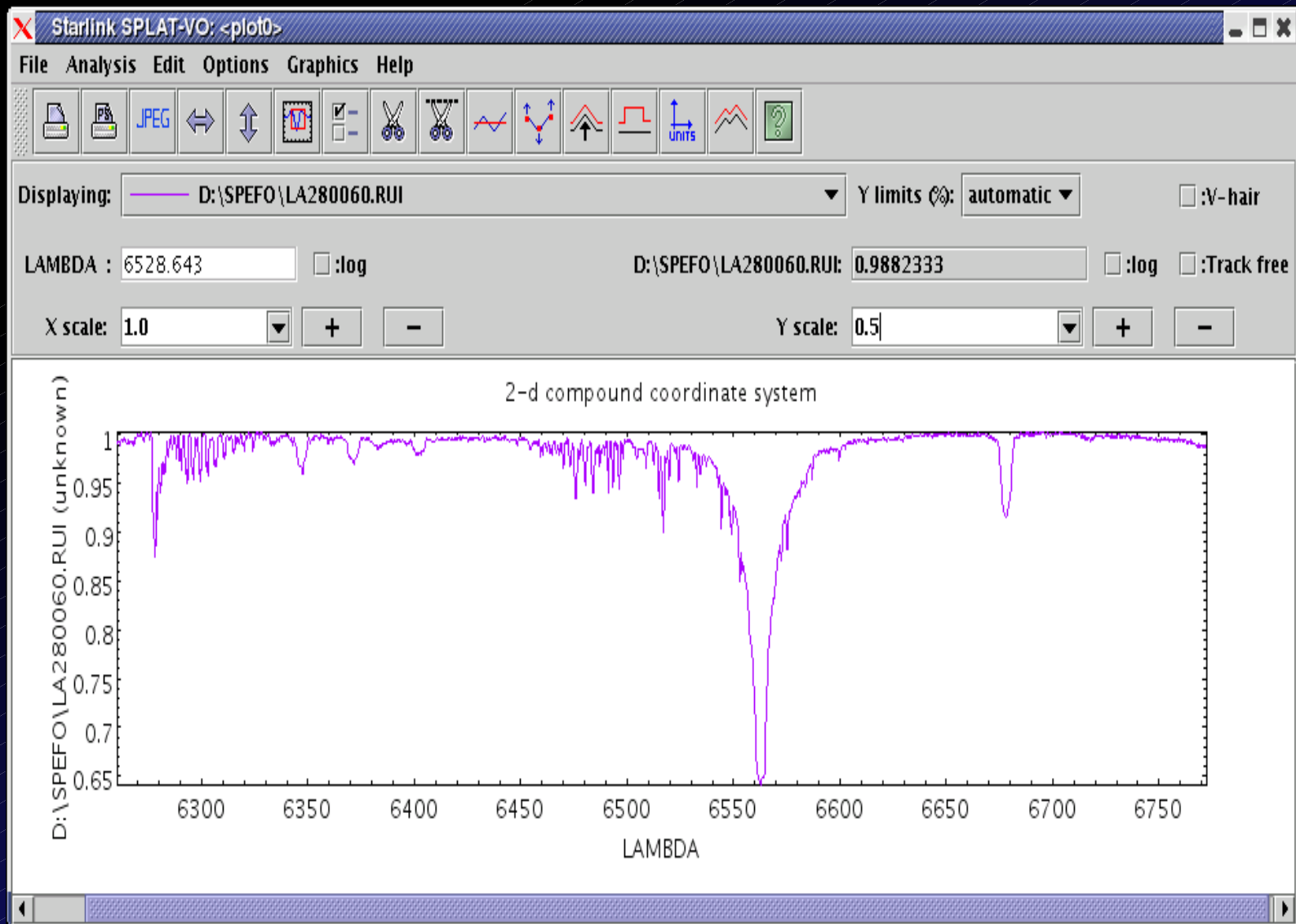
ViSiVio

VOSED

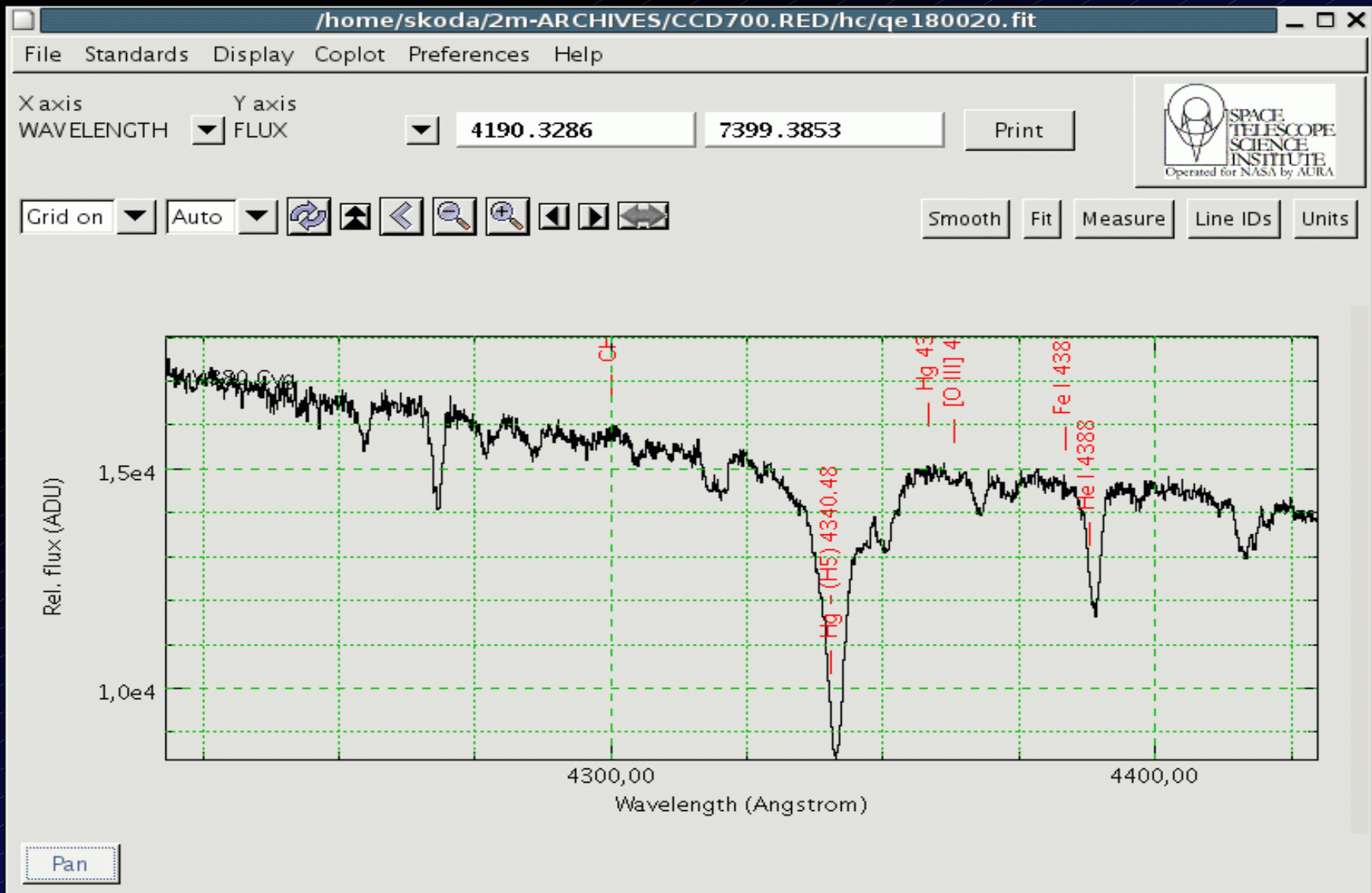
Sextractor – WESIX (Web Enabled Source Identification with Cross Matching)

PLASTIC HUB, Treeview + Period04 (since 18.9.08)

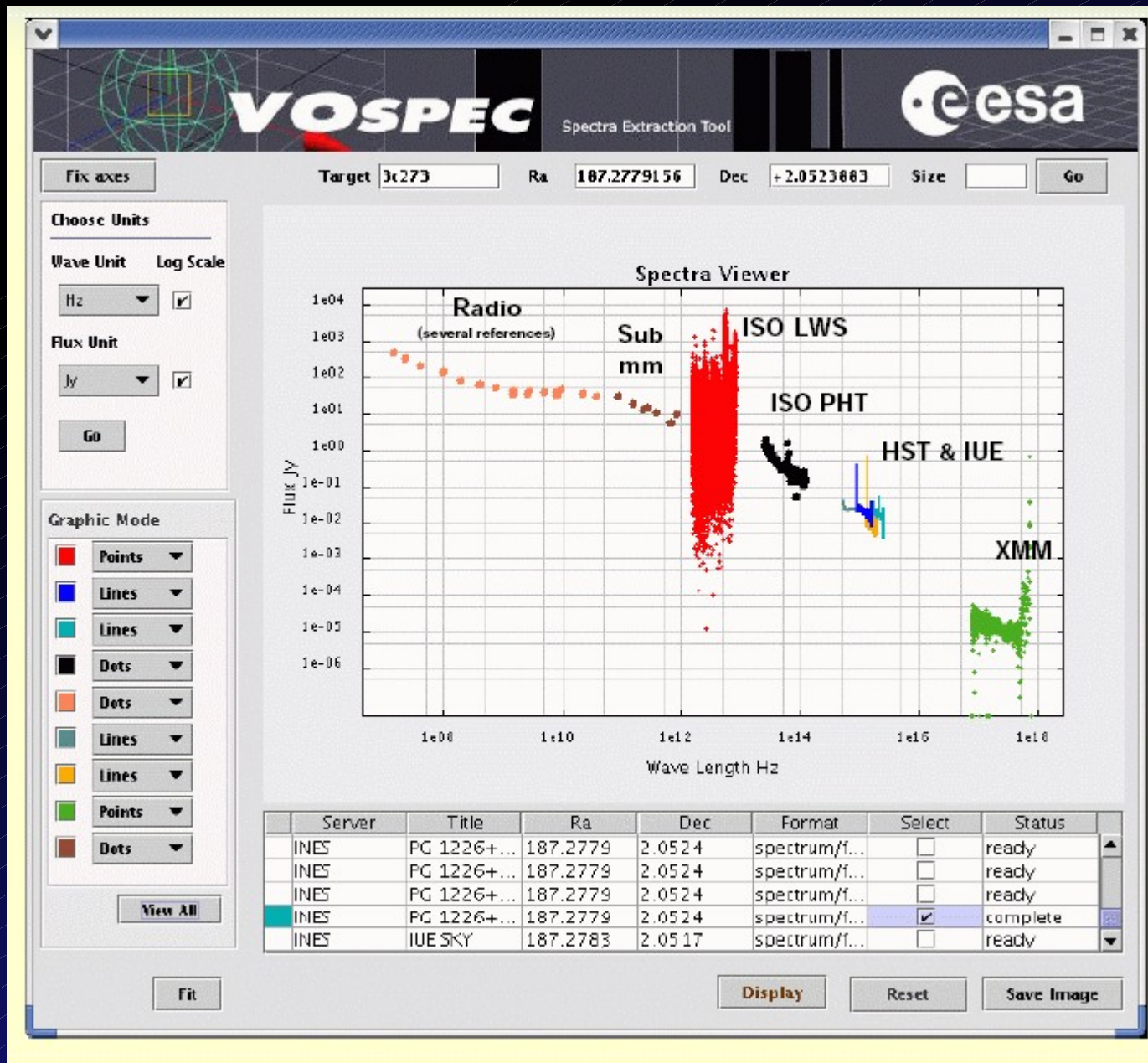
SPLAT-VO (Starlink, JAC)



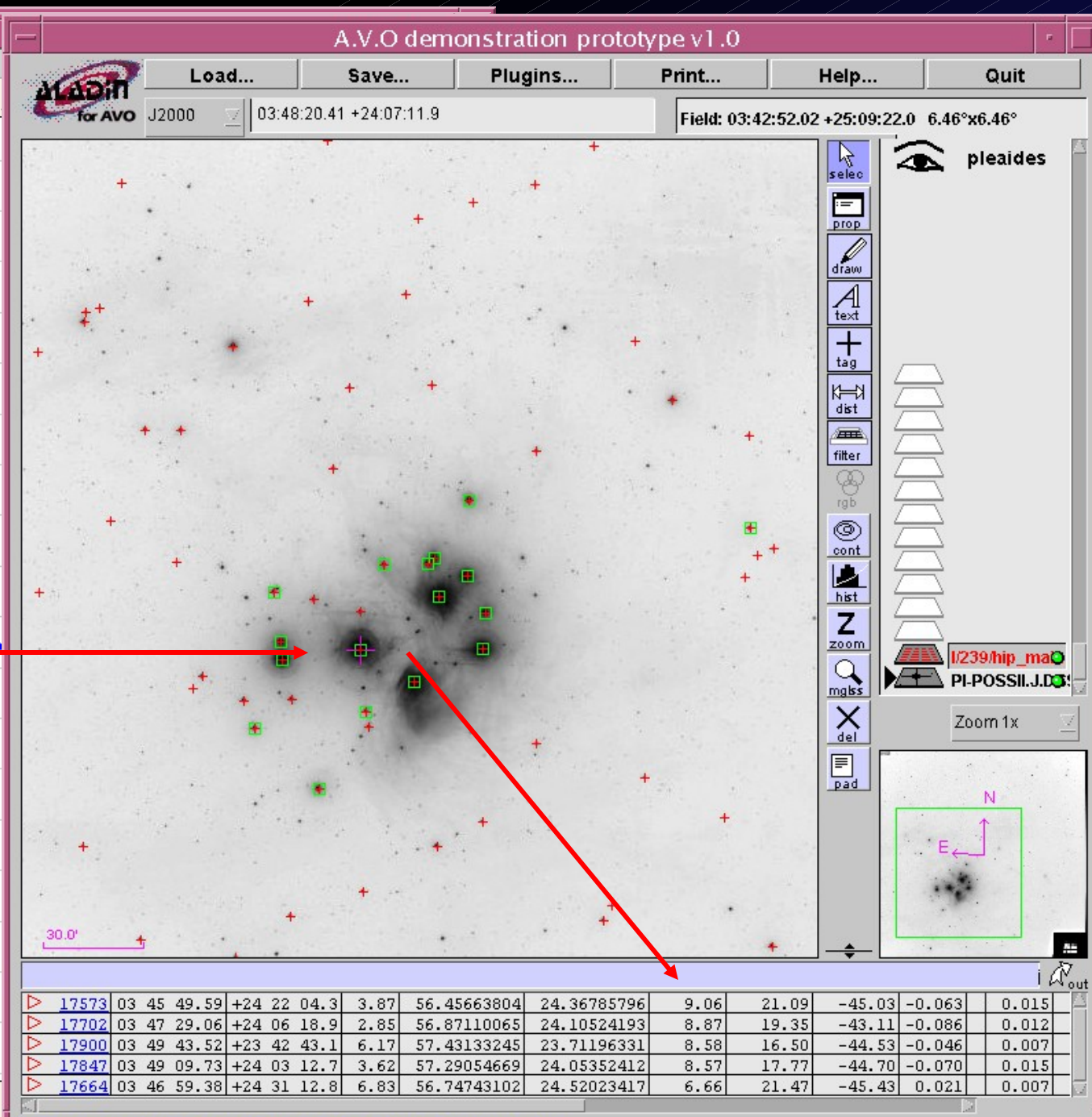
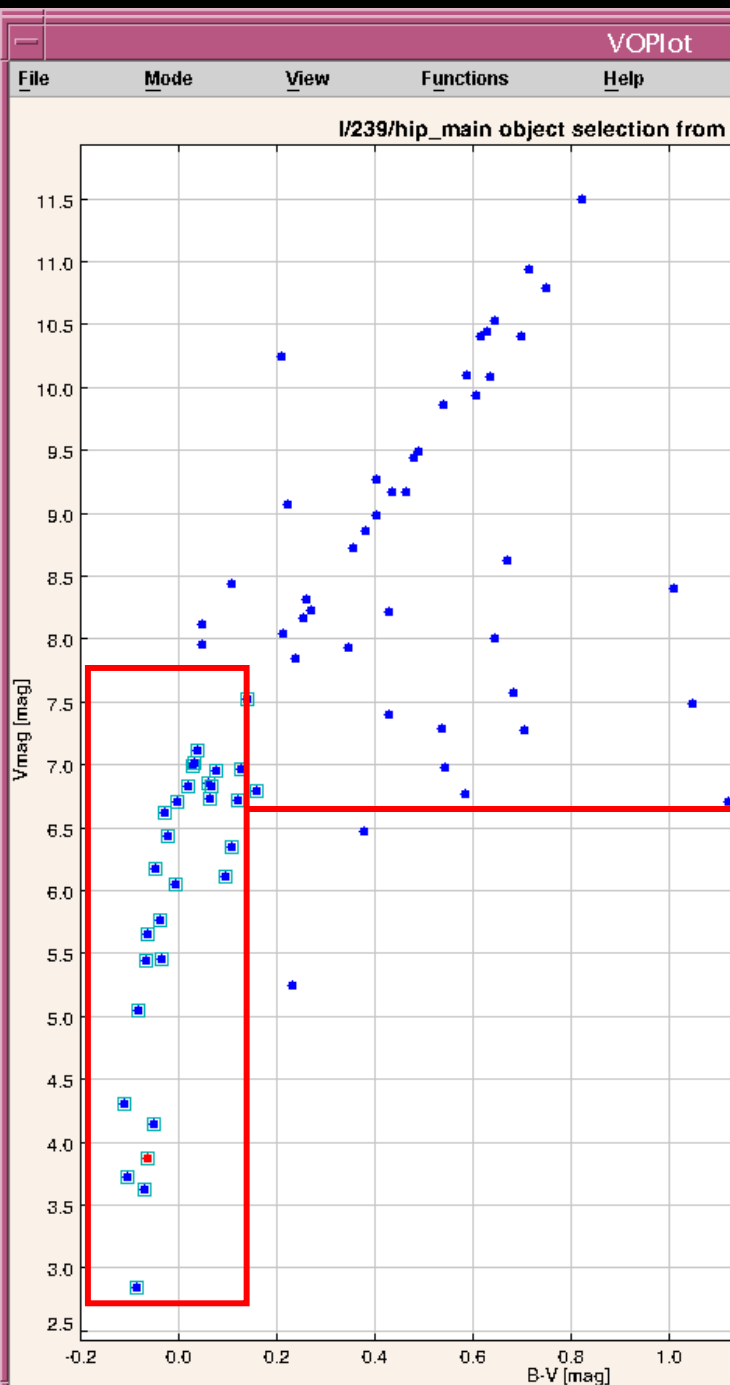
SpecView (STSci)



VOSpec (ESAC)

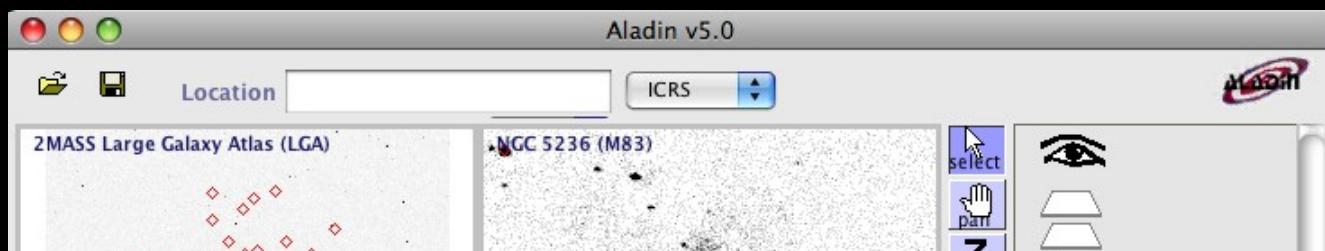


Colour-magnitude diagram

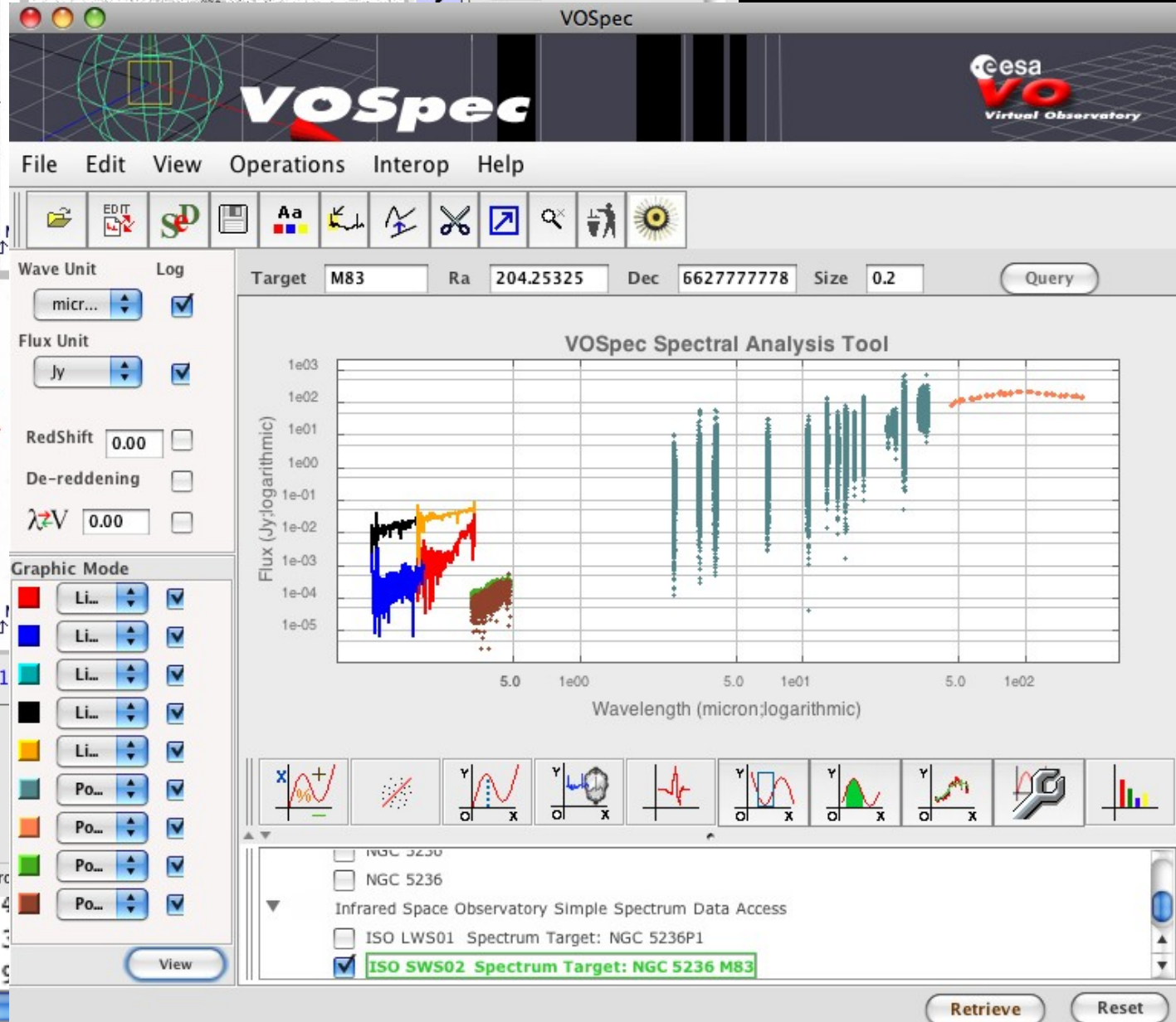


You are currently in Select mode.

CDS - ESO - AstroGrid - ST-ECF - UMAN/Jodrell Bank - CNRS/DR01 - VO-India - STScI



Tools



CIELO VO - line catalogue SLAP

SLAP Viewer Copyright ESAC, Spain

Server Selector

SLAP Services

- IASD
- LERMA
- NIST ATOMIC SPECTRA
- CIELO SLAP

Molecular line databases

<http://esavo02:8080/cieloslapToolKit/cieloslap.jsp?>

Select

Range of Search (m)

Wavelength Start Wavelength End

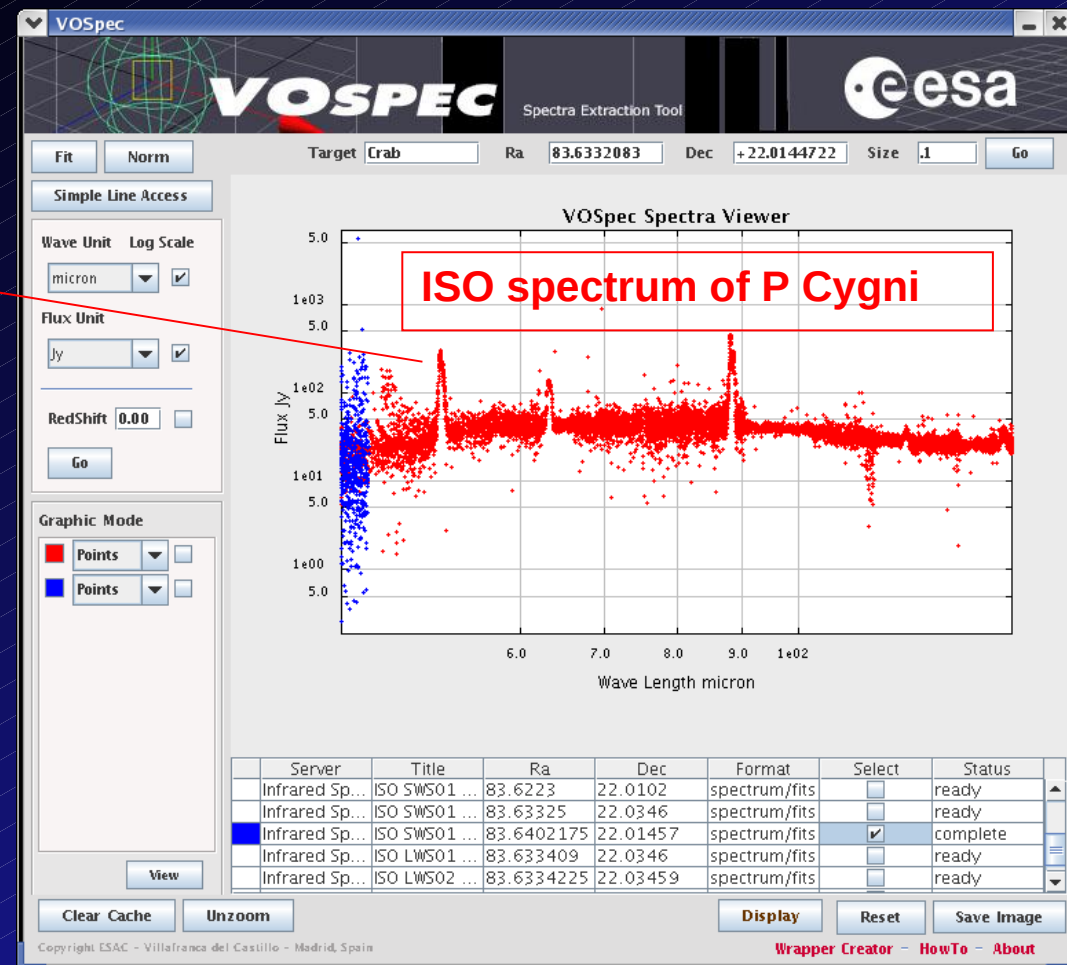
Reset

Slap Services Output

CIELO SLAP

Idm:Line.wavelength	Idm:Source...	Source.co...	Source.co...	Idm:Li...	Idm:...	Id...	Id...	Idm:...	I...
1.8627e-09	NGC1068	40.66963	-0.01328	1s_3p	1s2	1P1	1S0	OVII	...
1.7768e-09	NGC1068	40.66963	-0.01328	1s_4p	1s2	1P1	1S0	OVII	...
1.89671e-09	NGC1068	40.66963	-0.01328	2p	1s	2...	2...	OVIII	...
2.47793e-09	NGC1068	40.66963	-0.01328	2p	1s	2...	2...	NVII	...
2.21012e-09	NGC1068	40.66963	-0.01328	1s_2s	1s2	3S1	1S0	OVII	...
2.1602e-09	NGC1068	40.66963	-0.01328	1s_2p	1s2	1P1	1S0	OVII	...
2.18071e-09	NGC1068	40.66963	-0.01328	1s_2p	1s2	3P1	1S0	OVII	...
2.1621e-09	NGC1068	40.66963	-0.01328	1s_2p	1s2	3P1	1S0	OVII	...

Close



(IVOA Line Data Model: Dubernet, Osuna et al., in preparation)
(Simple Line Access Protocol: Salgado et al., in preparation)

VO for Atomic and Molecular Data

VAMDC (06/2009-12/2012 FP7)

13 organizations

Virtual Atomic and Molecular Data Centre

VO principles (web services, integration, registry, SAMP,
VODesktop, TOPCAT, VOSpec)

(includes VALD extractor, NIST)

extended citation system (all providers acknowledged)

Theory VO (TVO)

Methods of VO (parameters in DB, SQL...) for study of results of simulations , catalogues of simulated objects like SDSS...(PCA)

Browsing of simulation space along different axes – parameters, regions...

Evolutionary tracks, Photo Dissociation Regions

Formation of artificial galaxies, clusters – N body models (MR - 10 billions, 25TB)

Theoretical Spectra

Millennium Run

10^{10} particles

Several Gpc to
10 kpc

Cube 2 billion ly

One month MPSSC
25 TB

Evolution of 20 mil
galaxies

Evolution merger tree

Millennium Run
10,077,696,000 particles

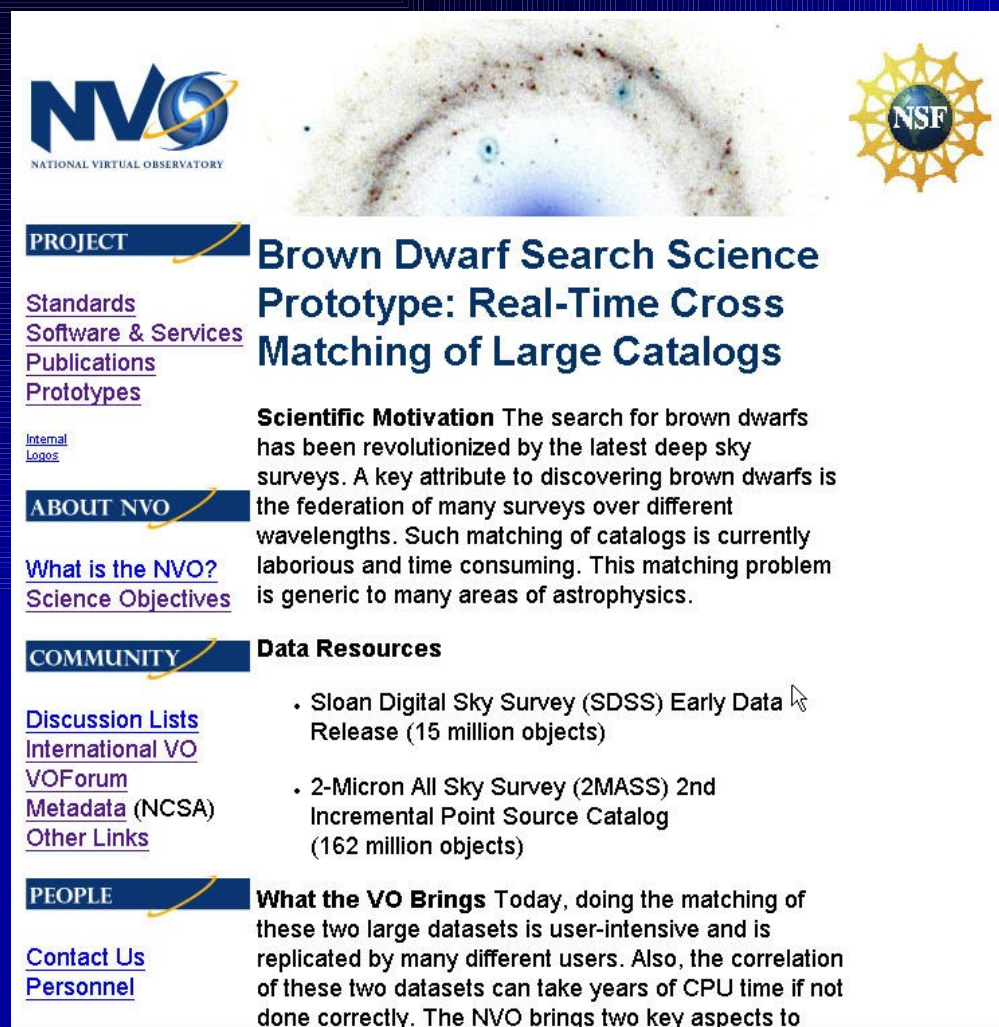
Spring et al. (2004)
Max-Planck-Institut für
Astrophysik



Examples of VO Science

- Discovery of obscured QSO (Padovani)
- Brown dwarfs (about 20 candidates)
- Brightest (WD?) (Caballero et al.) Albus-1
- Widest CPM binaries
- AGB to PN (100 new to 200 known using VO)
- SED (Spectrum Energy Distribution)
- Bolometric magnitude
- VOEvent – robotic telescopes (GRB, transits,)
- Outreach , Education (MS WWT, GoogleSky)

BDs discovered using VO



The screenshot shows the NVO Project website. At the top left is the NVO logo (National Virtual Observatory) and at the top right is the NSF logo. The main heading is "PROJECT Brown Dwarf Search Science Prototype: Real-Time Cross Matching of Large Catalogs". Below this, there are links for "Standards", "Software & Services", "Publications", and "Prototypes". A section titled "ABOUT NVO" contains links for "What is the NVO?", "Science Objectives", and "COMMUNITY". The "COMMUNITY" section lists "Data Resources" including "Sloan Digital Sky Survey (SDSS) Early Data Release (15 million objects)" and "2-Micron All Sky Survey (2MASS) 2nd Incremental Point Source Catalog (162 million objects)". A "PEOPLE" section at the bottom has links for "Contact Us" and "Personnel". A large astronomical image of a spiral galaxy is in the background.

PROJECT **Brown Dwarf Search Science Prototype: Real-Time Cross Matching of Large Catalogs**

[Standards](#)
[Software & Services](#)
[Publications](#)
[Prototypes](#)

[Internal Logos](#)

ABOUT NVO

[What is the NVO?](#)
[Science Objectives](#)

COMMUNITY

[Discussion Lists](#)
[International VO](#)
[VOForum](#)
[Metadata \(NCSA\)](#)
[Other Links](#)

PEOPLE

[Contact Us](#)
[Personnel](#)

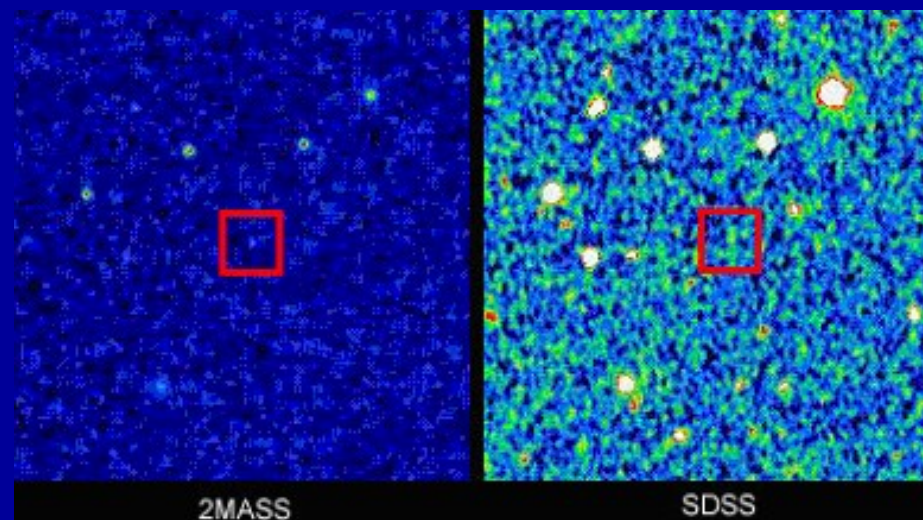
Scientific Motivation The search for brown dwarfs has been revolutionized by the latest deep sky surveys. A key attribute to discovering brown dwarfs is the federation of many surveys over different wavelengths. Such matching of catalogs is currently laborious and time consuming. This matching problem is generic to many areas of astrophysics.

Data Resources

- Sloan Digital Sky Survey (SDSS) Early Data Release (15 million objects)
- 2-Micron All Sky Survey (2MASS) 2nd Incremental Point Source Catalog (162 million objects)

What the VO Brings Today, doing the matching of these two large datasets is user-intensive and is replicated by many different users. Also, the correlation of these two datasets can take years of CPU time if not done correctly. The NVO brings two key aspects to

- **Filtering criteria:** z & J -only detections with $z - J > 2.75$
- *SDSS: 15M obj.*
- *2MASS: 160M obj.*
- *300000 objects in common.*

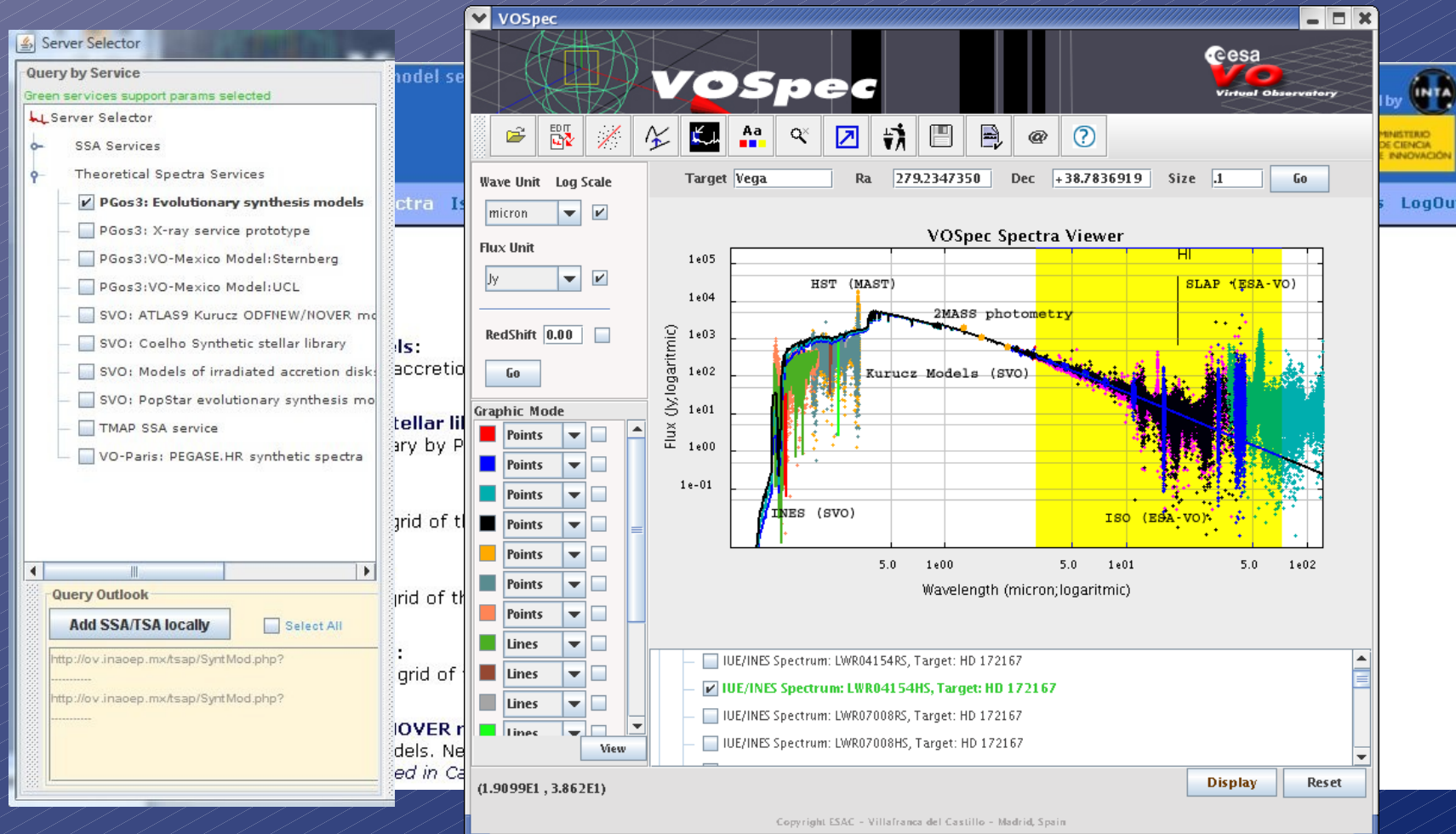


✓ *However, systematic searches using a VO methodology have not been performed so far.*

Archives, Theory, VO-Science, DataMining, E&O

Simple Spectral Access Protocol V1.04

Appendix A: Theoretical Spectral Access Use Case



Archives, Theory, VO-Science, DataMining, E&O

- Easy access to multi- λ info.
- Easy access to models (S3).

Astronomy & Astrophysics manuscript no. Synth VO PR1 ref format
August 2, 2008

VOSA: Virtual Observatory SED Analyzer.

An application to the Collinder 69 open cluster

A. Bayo^{1,2}, C. Rodrigo^{1,2}, D. Barrado y Navascués^{1,2}, E. Solano^{1,2}, R. Gutiérrez^{1,2}, M.
Morales-Calderón¹, and F. Allard³

Astronomy & Astrophysics manuscript no. 09595
April 14, 2008

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Young stars and brown dwarfs surrounding Alnilam (ϵ Ori) and Mintaka (δ Ori)

J. A. Caballero^{1,2} and E. Solano³

¹ Dpto. de Astrofísica y Ciencias de la Atmósfera, Facultad de Física, Universidad Complutense de Madrid, E-28040 Madrid, Spain

² Max-Planck-Institut für Astronomie, Königstuhl 17, D-69117 Heidelberg, Germany

³ Laboratorio de Astrofísica Espacial y Física Fundamental, INSA, Apdo. 78, E-28691 Villanueva de la Cañada, Madrid, Spain

Archives, Theory, VO-Science, DataMining, E&O

- Discovery of exotic objects. Access to large databases
 - 2MASS, SDSS, UKIDSS, IPHAS,...

Astronomy & Astrophysics manuscript no. aaVLM-iphas-v1
October 22, 2008

© ESO 2008

An IPHAS-based search for accreting very low-mass objects using VO tools

L. Valdivielso¹, E. L. Martín^{1,2}, H. Bouy^{1,3}, E. Solano^{4,5}, J. E. Drew^{6,7}, R.
and Vink, J. S.¹¹

Albus 1: A very bright white dwarf candidate

José Antonio Caballero^{1,2}

Max-Planck-Institut für Astronomie, Königstuhl 17, D-69117 Heidelberg, Germany

caballero@mpia.de

DRAFT VERSION NOVEMBER 6, 2009
Preprint typeset using L^AT_EX style emulateapj v. 05/04/06

THE FIRST L SUBDWARF IDENTIFIED IN THE UKIDSS INFRARED DEEP SKY SURVEY

N. LODIEU^{1,2}, M. R. ZAPATERO OSORIO³, E. L. MARTÍN³, E. SOLANO³, M. ABERASTURI³

Draft version November 6, 2009

Archives, Theory, VO-Science, DataMining, E&O

Theoretical model services

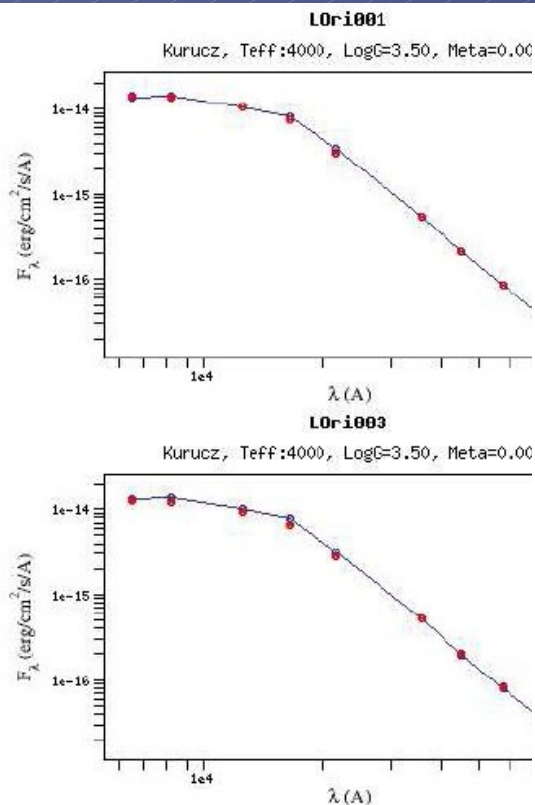


VOSA: VO Sed Analyzer

VO SED Analyzer

Services: VOSA Filters TSAP S3if

Object	Model	T_{eff}	LogG	Metallicity	χ^2
LO-i001	Kurucz	4000	3.50	0.00	1.71e
LO-i002	NextGen	3800	4.5	0	1.44e
LO-i003	Kurucz	4000	3.50	0.00	5.05e
LO-i004	Kurucz	3750	4.00	0.00	2.82e
LO-i005	NextGen	4000	4.0	0	9.20e
LO-i006	Kurucz	4000	3.50	0.00	3.36e
LO-i007	Kurucz	4000	4.50	0.00	2.49e
LO-i008	Kurucz	4000	3.50	0.00	4.43e
LO-i009	NextGen	4000	3.5	0	6.22e
LO-i010	NextGen	4200	4.0	0	1.87e
LO-i011	NextGen	3900	4.5	0	1.20e
LO-i012	NextGen	4000	4.5	0	6.58e
LO-i013	NextGen	3700	4.5	0	3.50e
LO-i014	Kurucz	4000	4.50	0.00	4.70e
LO-i015	Kurucz	4000	3.50	0.00	8.73e
LO-i016	Kurucz	3750	4.50	0.00	5.53e
LO-i017	NextGen	4200	4.0	0	7.58e
LO-i018	Kurucz	3750	3.50	0.00	4.31e
LO-i019	Kurucz	3750	3.50	0.00	1.90e
LO-i020	NextGen	3800	4.5	0	2.98e
LO-i021	Kurucz	4000	3.50	0.00	3.08e
LO-i022	Kurucz	3750	4.00	0.00	1.76e
LO-i023	NextGen	4000	4.5	0	2.35e
LO-i024	Kurucz	3750	3.50	0.00	2.22e
LO-i025	NextGen	3700	4.5	0	1.37e
LO-i026	NextGen	3700	4.5	0	4.81e
LO-i027	NextGen	4000	4.5	0	2.35e
LO-i028	Kurucz	3750	4.00	0.00	1.28e
LO-i029	NextGen	3100	4.5	0	7.28e
LO-i030	NextGen	3700	4.5	0	2.15e
LO-i031	NextGen	3800	4.5	0	2.77e
LO-i032	NextGen	3700	4.5	0	1.70e
LO-i033	NextGen	3700	4.5	0	7.12e
LO-i034	NextGen	3000	4.0	0	1.77e
LO-i035	NextGen	3700	4.5	0	1.81e



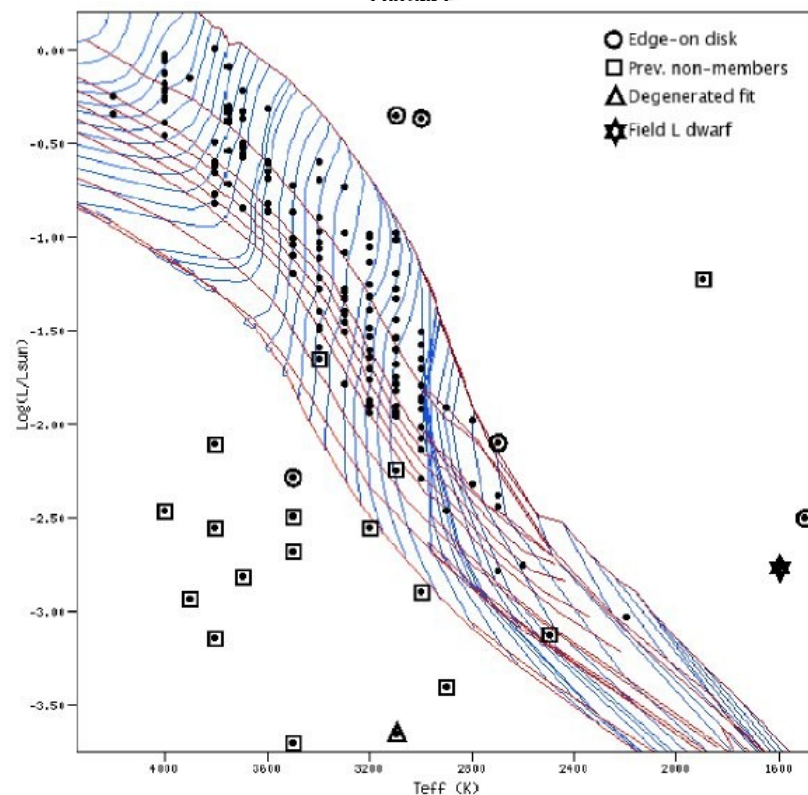
Astronomy & Astrophysics manuscript no. Synth VO PR1 ref format

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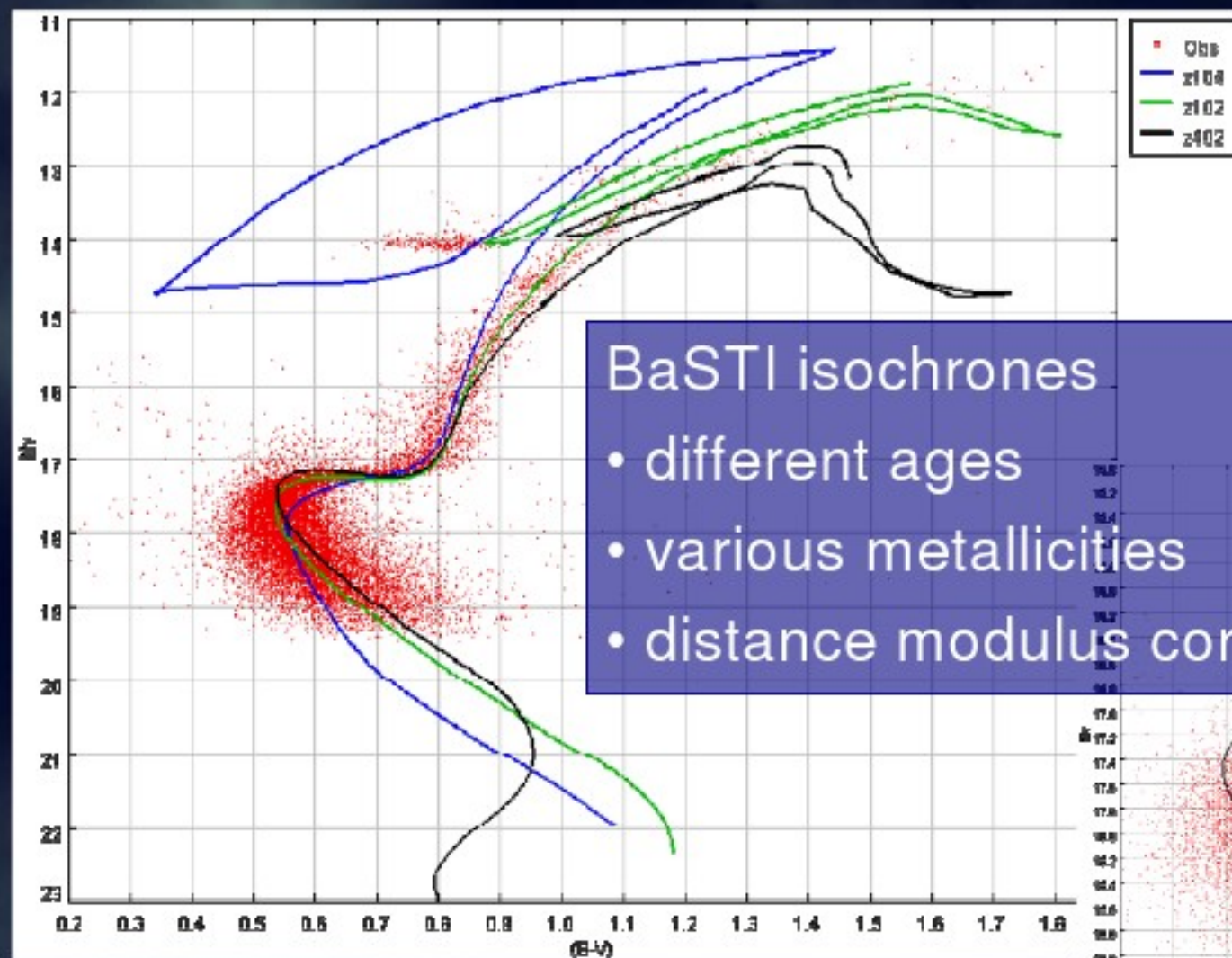
VOSA: Virtual Observatory SED Analyzer.

An application to the Collinder 69 open cluster

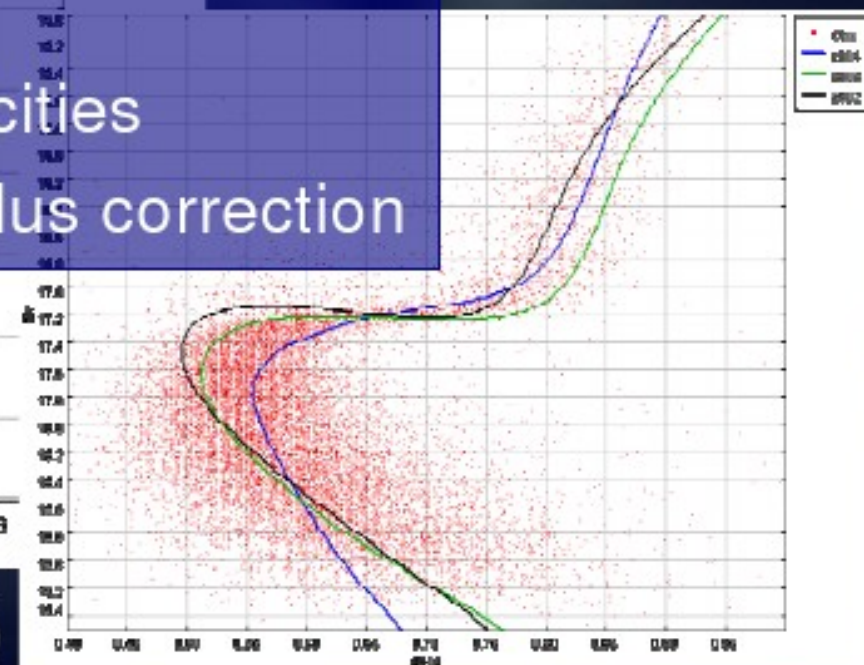
A. Bayo^{1,2}, C. Rodrigo^{1,2}, D. Barrado y Navascués^{1,2}, E. Solano^{1,2}, R. Gutiérrez^{1,2}, M.



BaSTI Isochrones



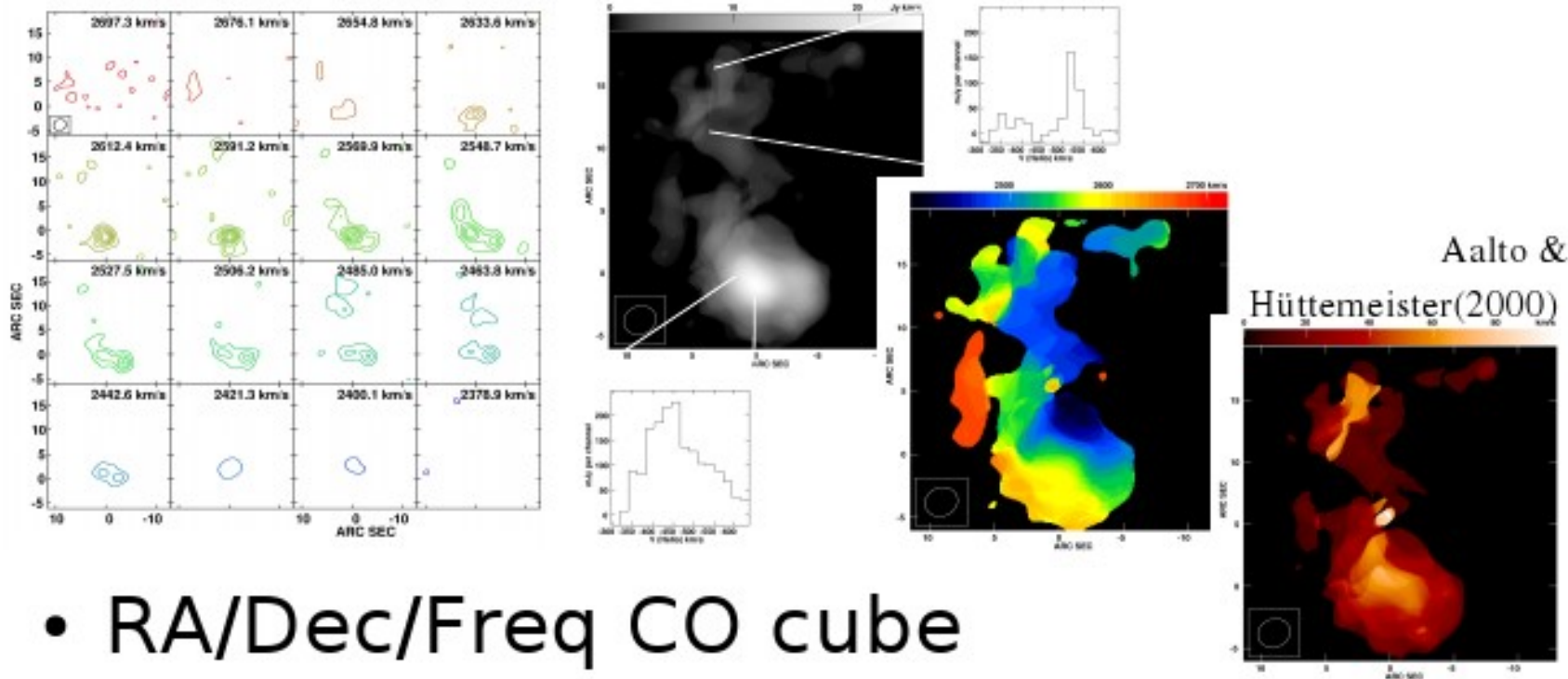
step 1
metallicity



$z = 0.01$ (α -enh) ; 0.008 (scaled solar)



ALMA/IRAM use case



- RA/Dec/Freq CO cube
 - Convert to velocity (LSR, radio convention)
 - Cutouts, simple squashes - VO tools?
 - Smoothed spectra, moments with noise cut-off
 - Specialised server-side pipeline controlled via UWS

Objections to VO

Data quality – garbage in - garbage out

How and whom to give credit ? (button)

embedded ivo:// data in ApJ

VO for dissemination only

technology for OPTICON, nextgen

Virtual science – VO technology

VO only for public data ! Proprietary ?

(data jealousy)

local archive - available data marked

Other VOs

Virtual Solar Observatory

Virtual Solar-Terrestrial Observatory

Virtual Magnetospheric Observatory

Virtual Space Physics Observatory

Virtual Meteor Observatory

Interest of climatology, meteorology

New branch of Science = e-Science

Democratization of Science

- Digital Divide (data access free, journals ?)
- International Council for Science (ICS UNO)
CODATA (Committee on Data for Science and Technology)
- OECD, UNESCO
- CASPAR

Cultural, Artistic and Scientific knowledge for Preservation, Access and Retrieval

Digital curation centers

ADS and VO (links to ivo://, metadata, ontologies – understanding, semantic web)

Archive importance: 5x IUE , 3x HST results from archives than PI articles

Effectivity – 50% of published data appears in Journals, links to data automatic ?

The Astronomer's Data Manifesto

at 26 IAU GA Prague SPS3

- (a) All significant tables, images, and spectra published in journals should appear in astronomical data centres.
- (b) All data obtained with publicly-funded observatories should, after appropriate pro-prietary periods, be placed in the public domain.
- (c) In any new major astronomical construction project, the data processing, storage, migration, and management requirements should be built in at an early stage of the project plan, and costed along with other parts of the project.
- (d) Astronomers in all countries should have the same access to astronomical data and information.
- (e) Legacy astronomical data can be valuable, and high-priority legacy data should be preserved and stored in digital form in the data centres.
- (f) The IAU should work with other international organisations to achieve our common goals and learn from our colleagues in other fields. ”

CZVO

VO-KOREL (web services)

parallel run of many jobs – more users

web service using VO Universal Worker Server

job control, queuing, jobs results polling

will be integrated in VODesktop

1D spectra cutout server (HEROS)

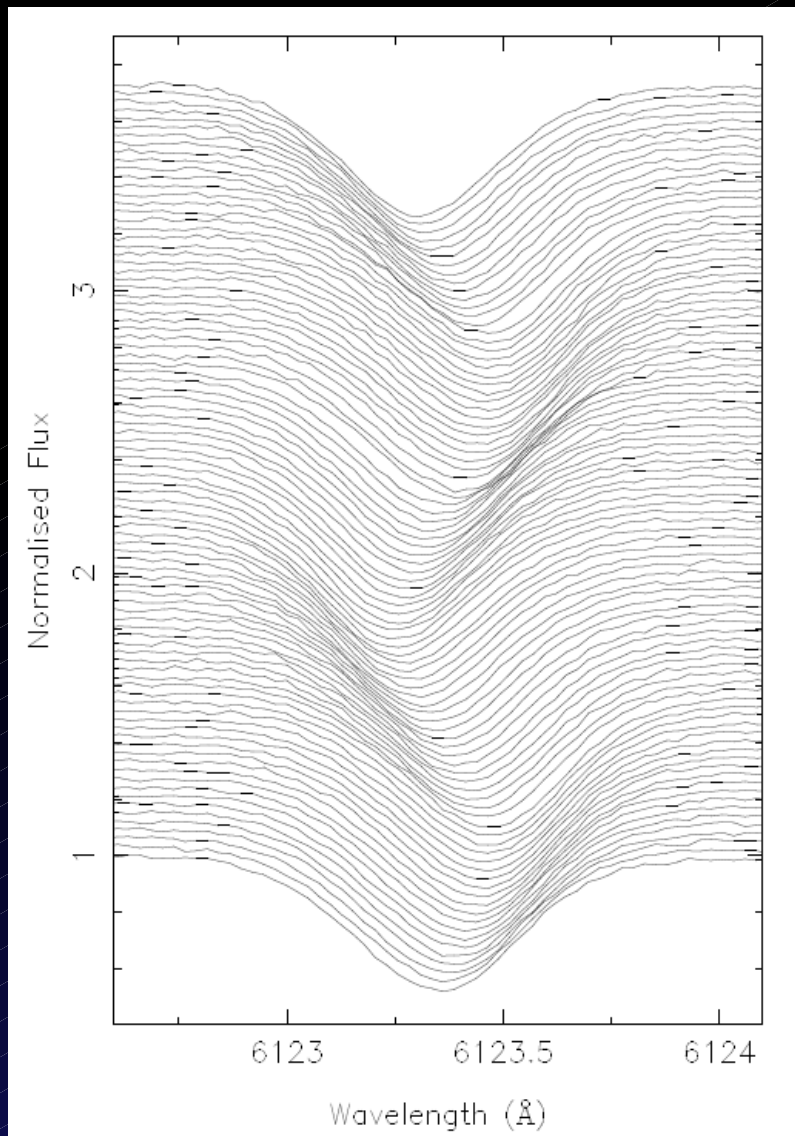
SSA access to 1D spectra + cutout of regions (lines)

need normalization, rebinning , convolution
(resolution) on server

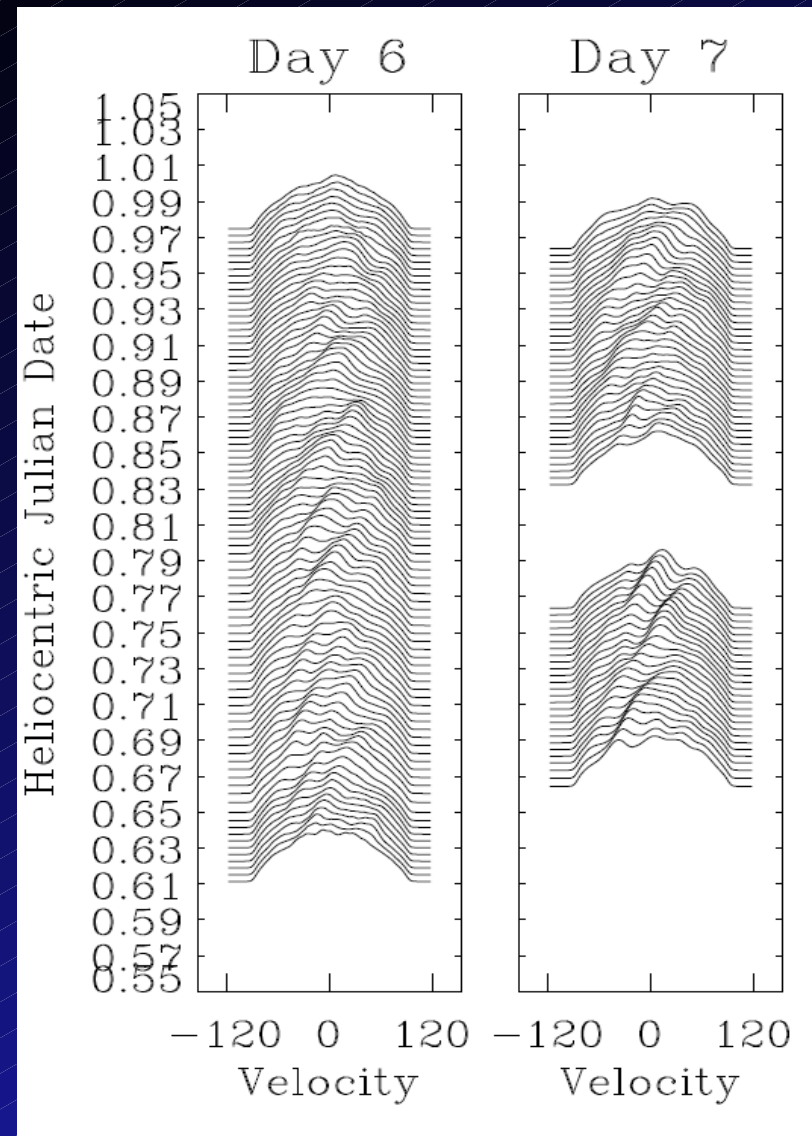
other 1D FITS spectra easy (Rozhen ?)

Data mining – chaos, NN

Measured Pulsations



Rho Pup – Aerts



Eps Cep - Aerts

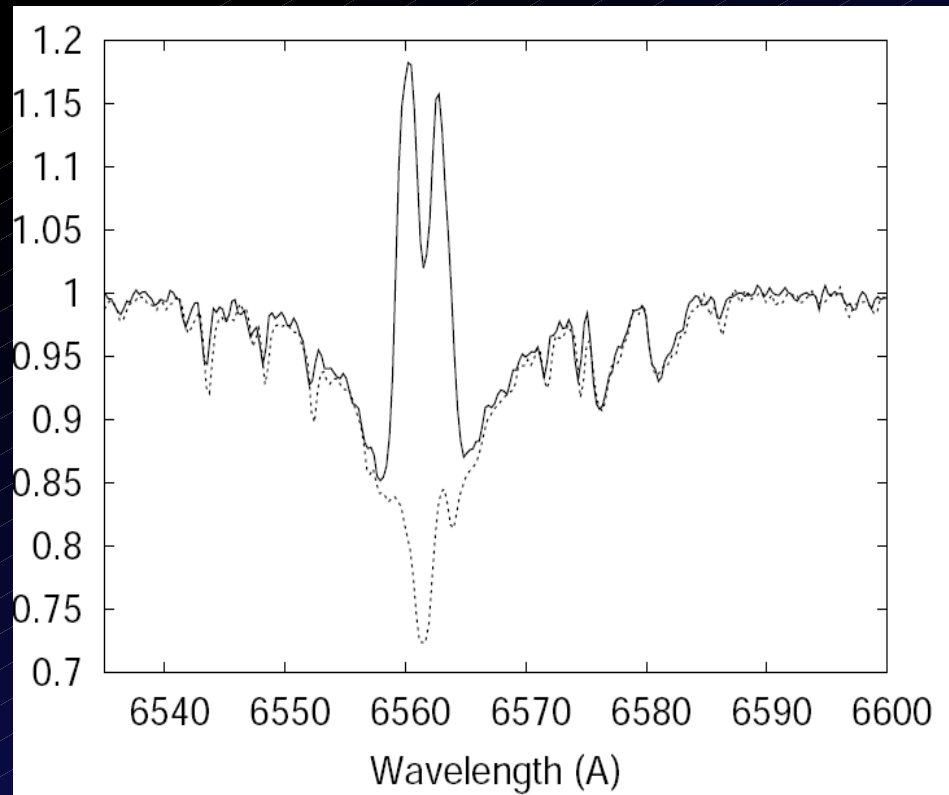
Changes of Line Profiles in Time

Blind comparison of different exposures

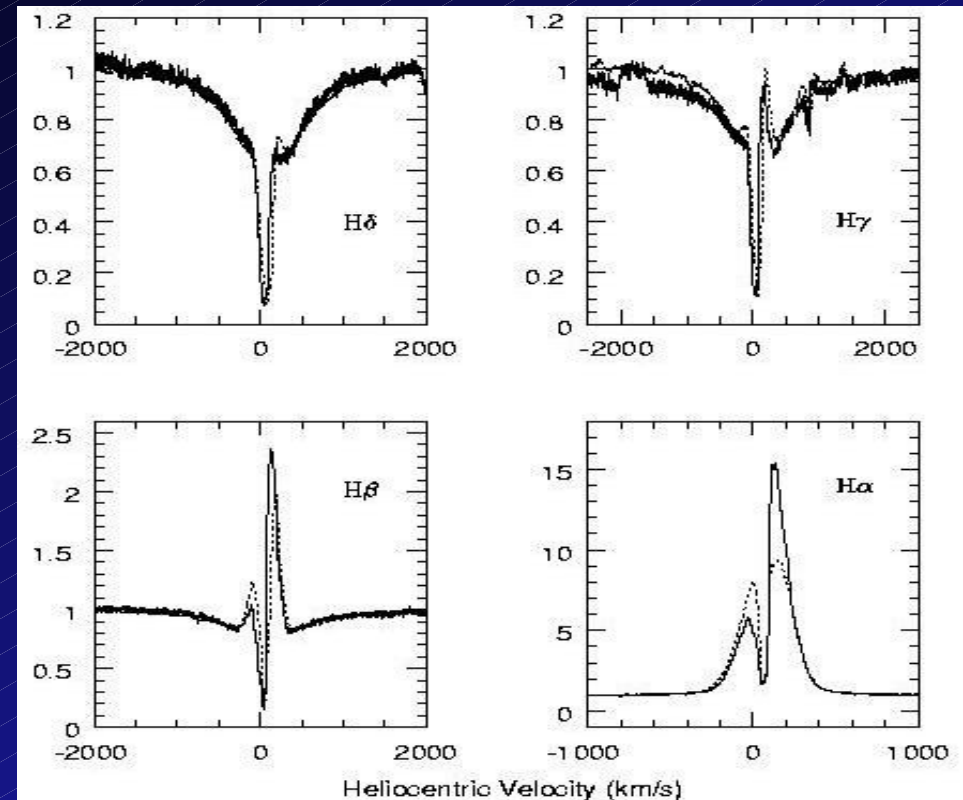
Emission/absorption, shell phases

Time evolution of object – mass transfer,

V/R variations



HD6226 : Slechta and Skoda 2004



Borges et al. 2008

Spectra postprocessing

- cutout service (ranges in SSAP)
- Rebinning (change of $\Delta\lambda$ in echelle)
- Instrument profile (de)convolution
- Broadening functions (rotation, limb dark)
- RV shift
 - on client side (memory, size)
 - on server side (Pleinpot WWW pipeline)

Killer spectral applications

Use VO to find all stars with emission in given line ($EW < 0$) – find the time when it was in em.

Use VO to get 1000 spectra of the given object cut out regions around given lines, plot the lines, make a gray dynamic spectrum folded in time

The same – search period, fold by period

Get the unknown line ID of piece of spectra from SLAP overplotted over SSA data

Create Light and RV curve for given period

Fit the grid of models (T_{eff} , $\log g$) to the observed spectrum – for many stars

VIRTUAL OBSERVATORY

